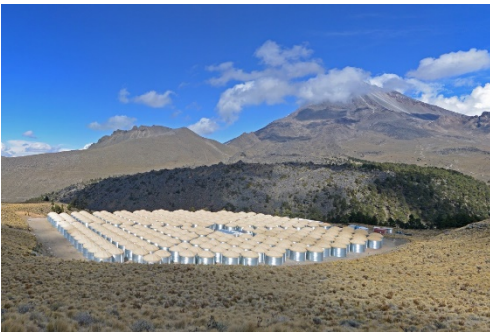




# Astrophysics in the NSF Physics Division

C. Denise Caldwell  
Division Director

With input from Jean Cottam, Jim Whitmore, Keith Dienes,  
Pedro Marronetti, Mark Coles, Allena Oppen, and Slava Lukin



Credit: HAWC



Credit: LIGO Lab

AAAC January 2018



# Perspectives on the Frontiers of Physics

**Controlling the Quantum World**– Electromagnetic radiation in the non-classical limit, Entanglement, Cavity QED, QIS, Optomechanics (Optical Physics; Quantum Information Science)

**Complex Systems and Collective Behavior** – Living cells, biological systems, ultracold fermions and bosons, quark-gluon liquid (Physics of Living Systems; Atomic and Molecular Dynamics; Nuclear Physics; Plasma Physics)

**Neutrinos and Beyond the Higgs** – Neutrino mass, new particles, unification of quantum mechanics and gravity, electron and neutron dipole moments (Particle Astrophysics; Gravitational Physics; Nuclear Physics; Precision Measurements; Elementary Particle Physics)

**Origin and Structure of the Universe** – Star formation and creation of the elements, dark matter and dark energy, modeling of black holes, gravitational waves, magnetic fields (Gravitational Physics; Nuclear Physics; Particle Astrophysics; Plasma Physics)

**Strongly-Interacting Systems**– QCD computations, quark structure of baryons, high-field laser-matter interactions, supernovae, strong gravity (Nuclear Physics; Gravitational Physics; Plasma Physics)

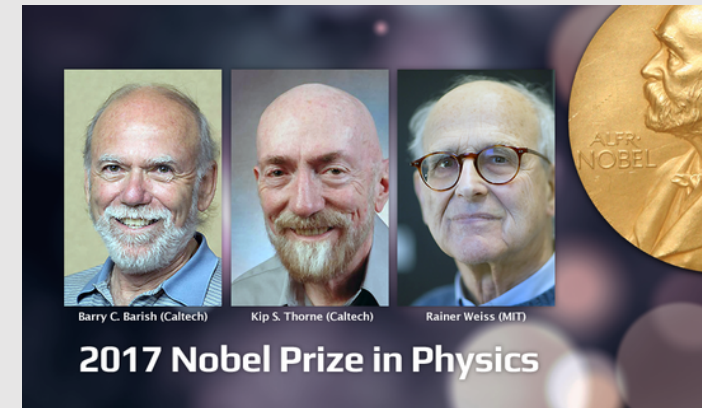
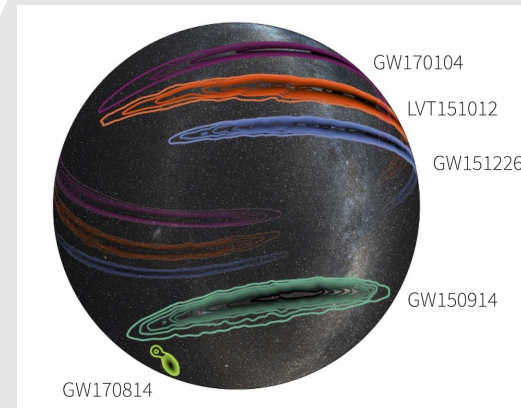


# LIGO

## An amazing year!

The 2<sup>nd</sup> Observational Run (O2) ended in August 2017 with an impressive collection of achievements

- Jan. 4: LIGO detects the 3<sup>rd</sup> binary black hole merger (BBH)
- June 8: LIGO detects the 4<sup>th</sup> BBH merger
- Aug. 14: First triple detection of a BBH merger (LIGO + Virgo)
- Aug. 17: First detection of a binary neutron star (BNS) merger
- Oct. 3: Weiss, Thorne and Barish win the 2017 Nobel Prize in Physics





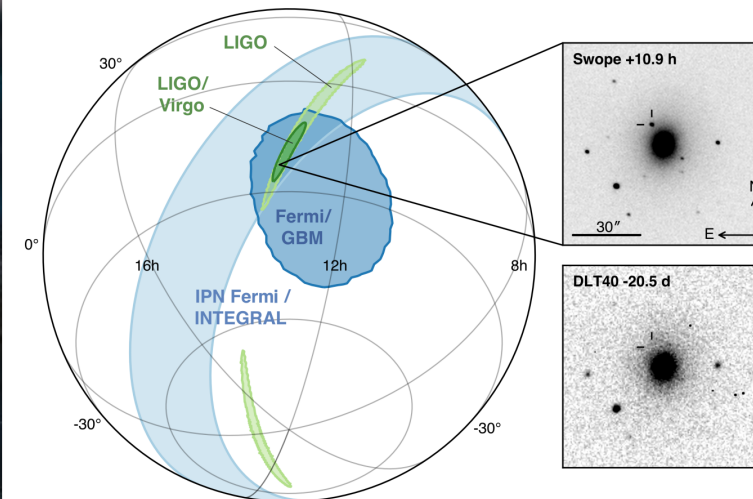
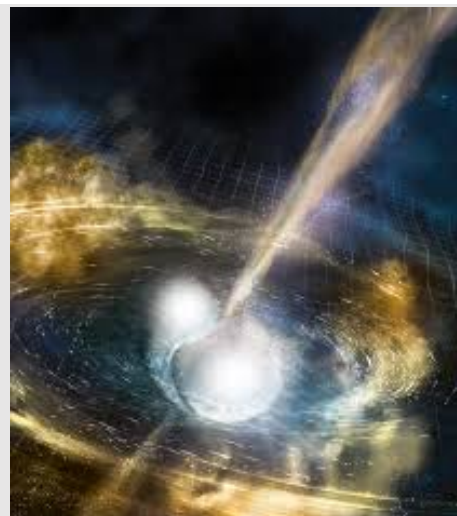


# LIGO

## The first binary neutron star merger detection

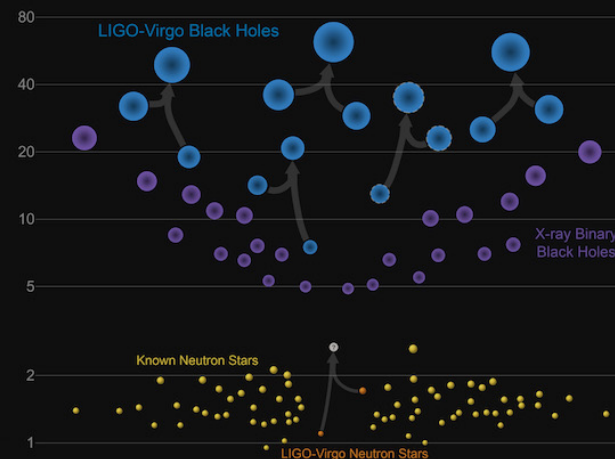
The joint observation of the first BNS merger by GW and EM observatories was a milestone in astronomical history, heralding the new era of Multimessenger astronomy

- Only 1.7 seconds after the GW detection, a gamma-ray burst (GRB) was detected by NASA Fermi satellite
- A massive EM follow up campaign determined that the event was located in the galaxy NGC 4993 (40 Mpc, 130 Mly)
- First direct evidence of the BNS merger as progenitors of short duration GRBs
- First clear observation of a kilonova (emissions of the radioactive decay of the BNS merger ejecta)
- New determination of the Hubble constant, independent of any form of the cosmic “distance ladder”



### Masses in the Stellar Graveyard

*in Solar Masses*





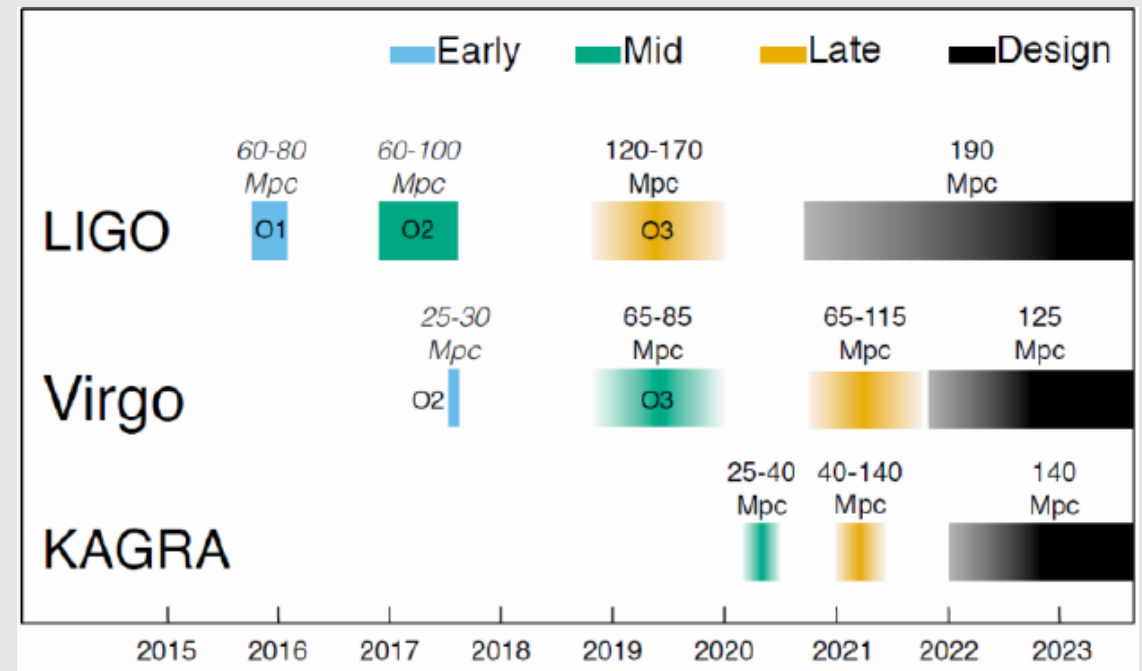


# LIGO

## Schedule for 2018

LIGO and Virgo are currently undergoing commissioning tasks to prepare for O3

- O3 Start in late Fall 2018
- Planned duration: 9 months
- Sensitivity at 2/3 of design level
- Virgo expected join in from the start
- LIGO India currently in land acquisition process. Ground breaking planned for mid 2018

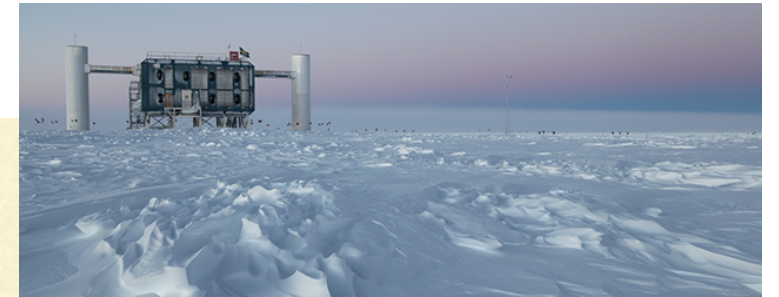




# IceCube UHE Neutrino on Sept. 22, 2017

Last year, had another “first” in multi-messenger discoveries, when MAGIC and Fermi-LAT measured VHE gamma rays coming from a direction consistent with a UHE neutrino event detected by the IceCube Neutrino Observatory in Antarctica.

## ICECUBE-170922A



GCN AMON notice:[https://gcn.gsfc.nasa.gov/notices\\_amon/50579430\\_130033.amon](https://gcn.gsfc.nasa.gov/notices_amon/50579430_130033.amon)

Date: 22 Sep, 2017

Time: 20:54:30.43 UTC

RA: 77.43 deg (-0.80 deg/+1.30 deg 90% PSF containment) J2000

Dec: 5.72 deg (-0.40 deg/+0.70 deg 90% PSF containment) J2000

E proxy of about 120 TeV

# EM activity follow-up of the IceCube telegram



## #1 Fermi-LAT detection of increased gamma-ray activity of TXS 0506+056, located inside the IceCube-170922A error region.

They found one source TXS 0506+056 – a flaring blazar, also known as 3FGL J0509.4+0541

## #2 AGILE confirmation of gamma-ray activity from the IceCube-170922A error region

They found a  $4\sigma$  significance centered less than 1 degree from both the IceCube centroid and the position of 3FGL j0509.4+0541

## #3 VLA Radio Observations of the blazar TXS 0506+056 associated with the IceCube-170922A neutrino event

They observed TXS 0506+056 over 4 epochs and frequency bands. They detect it significantly in all bands/epochs

## #4 First-time detection of VHE gamma rays by MAGIC from a direction consistent with the recent EHE neutrino event IceCube-170922A

They observed this source and found a  $5\sigma$  detection above 100 GeV after 12 hrs

## #5 IceCube-171106A: Swift observations

VERITAS, HAWC, ANTARES set upper limits More details to come in a paper.

SWIFT notes 5 X-ray sources in the IceCube error box!! 2 of which are luminous blazars.

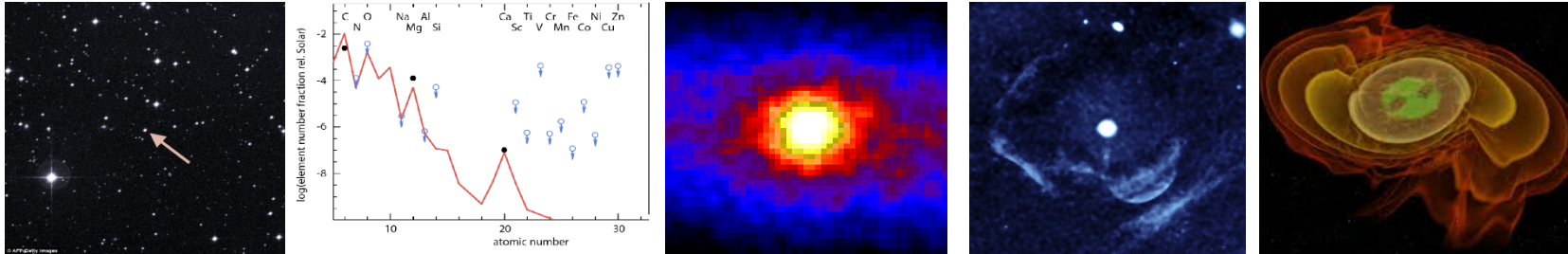




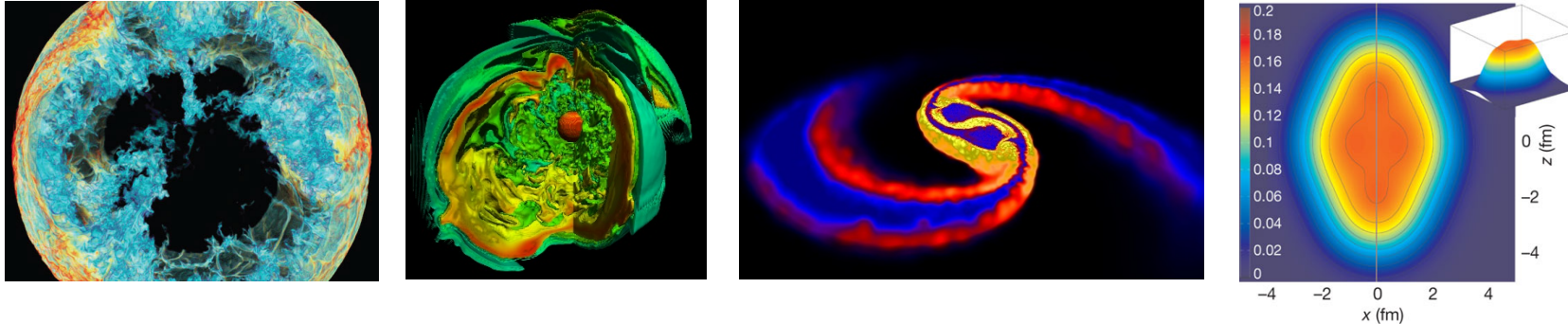
# Nuclear Astrophysics

A field at the interface of Astrophysics and Nuclear Physics

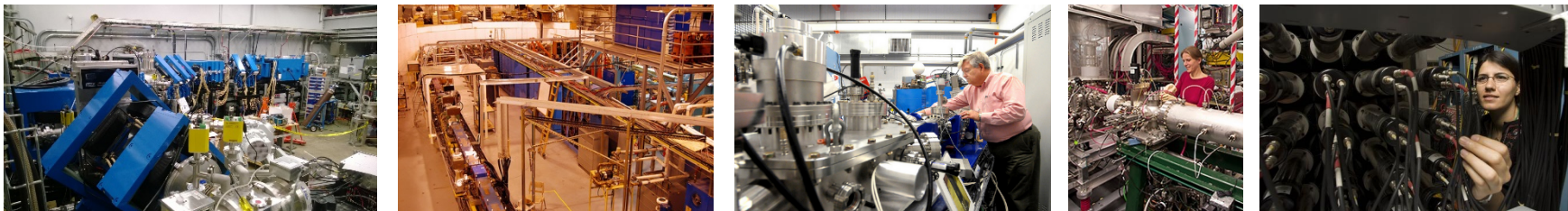
## ➤ Observational tools, signatures, and developments



## ➤ Large computational modeling for stellar and nuclear systems

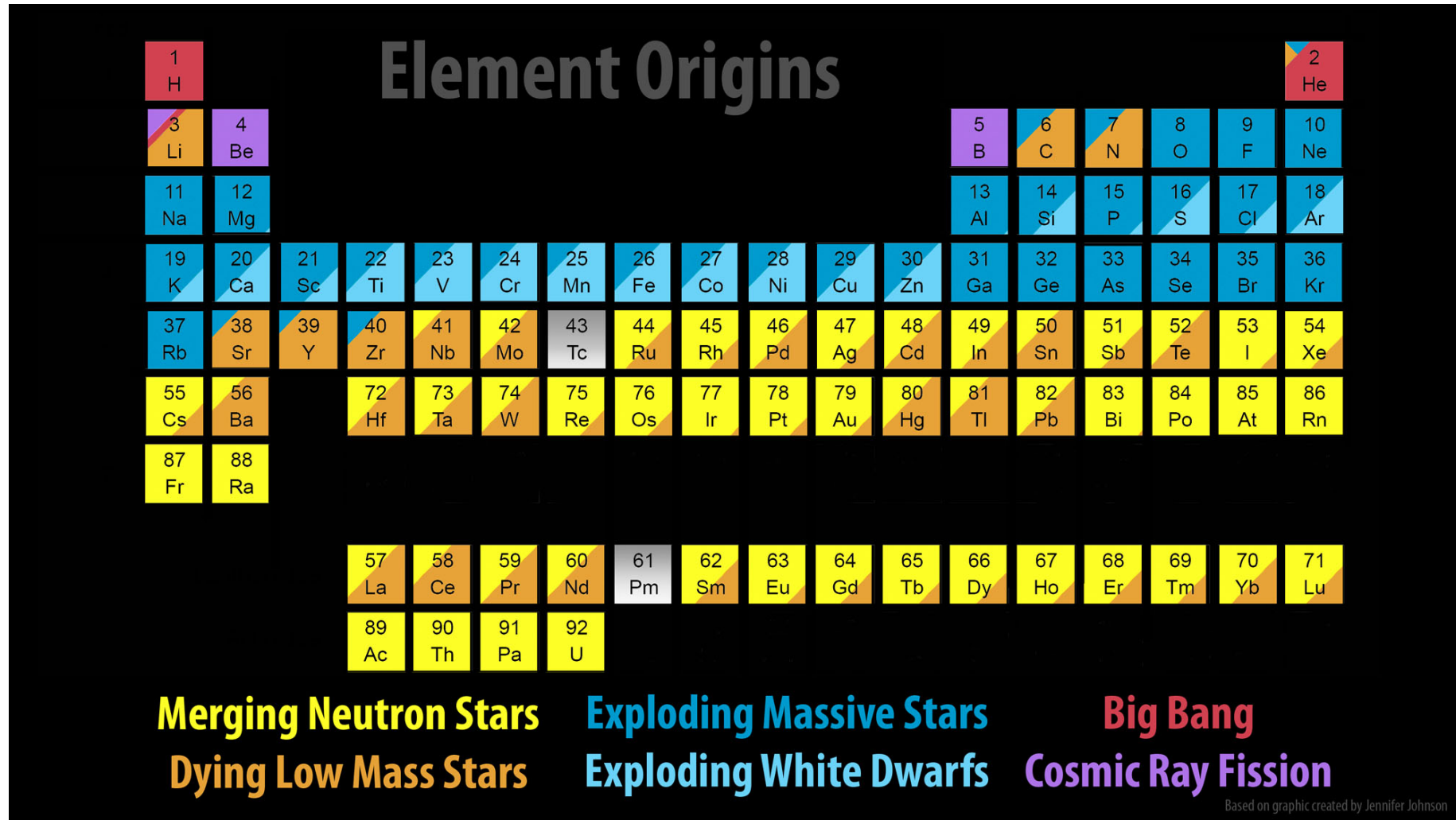


## ➤ Laboratory tools for experimental evidence





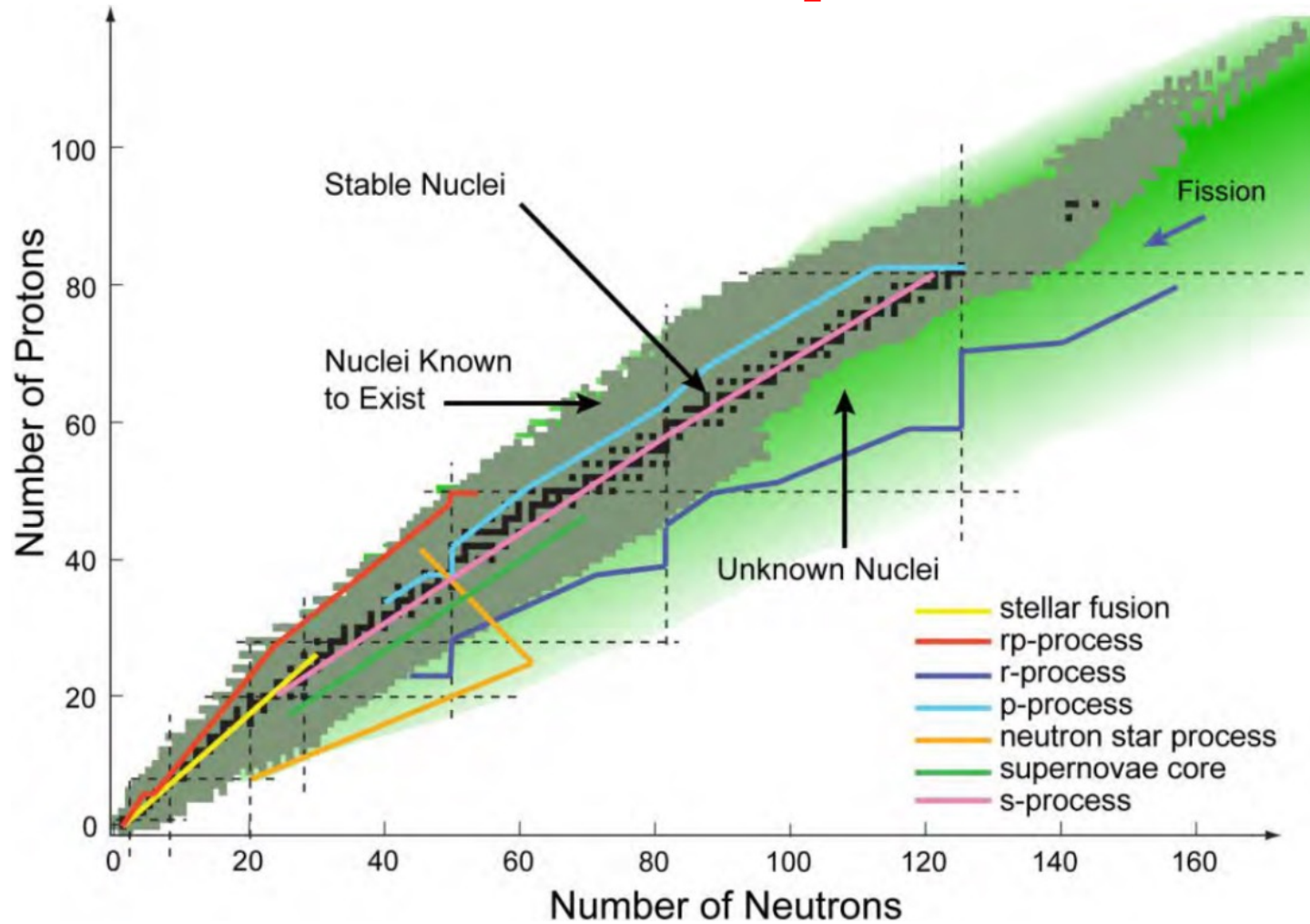
# Products of Neutron Star Merger



Credit: Jennifer Johnson/SDSS



Much more complicated!



Credit: NSCL/FRIB





# JINA-CEE LIVESTREAM

FRIDAY, 1 DECEMBER 2017

12:30 PM US/EST

[View the Recorded Event](#)



Talks and Panel Discussion:

The Impact of the LIGO/VIRGO Neutron Star Merger Discovery on  
Research in Nuclear Science and Nuclear Astrophysics

#GWNuclear

#GWNuclear

Program

Conclusions

INT-JINA Symposium INT-18-72R

## First multi-messenger observations of a neutron star merger and its implications for nuclear physics

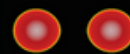
March 12 - 14, 2018

### Neutron Star Merger Dynamics

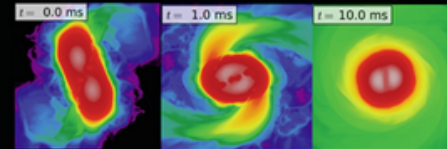
(General) Relativistic (Very) Heavy-Ion Collisions at  $\sim 100$  MeV/nucleon

Simulations: Rezzola et al (2013)

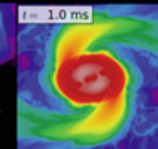
$t = -8.1$  ms



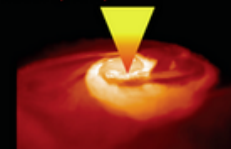
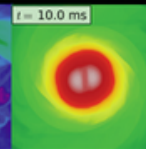
$t = 0.0$  ms



$t = 1.0$  ms



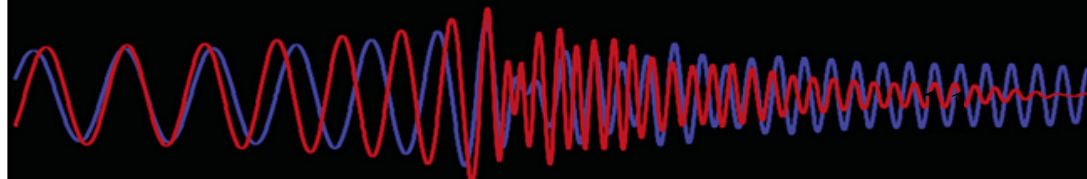
$t = 10.0$  ms



Inspiral:  
Gravitational waves,  
Tidal Effects

Merger:  
Disruption, NS oscillations, ejecta  
and r-process nucleosynthesis

Post Merger:  
GRBs, Afterglows, and  
Kilonova





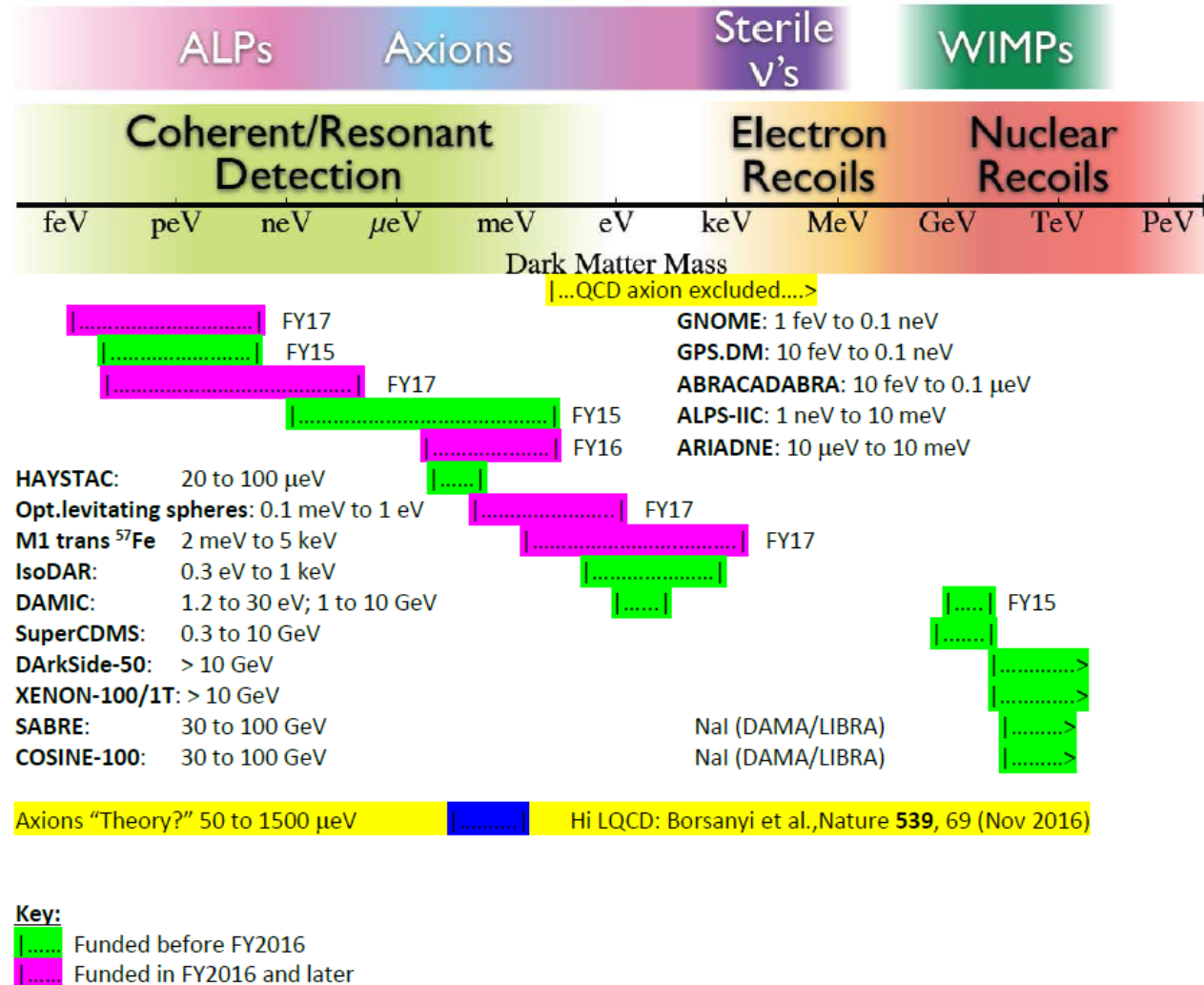
# Dark Matter Candidates (Update)

Dark Matter experiments funded by the NSF-Particle Astrophysics Program

**We need to search for low mass Dark Matter:**

**We are funding the following projects searching in the region for masses < 1 MeV**

**(Purple are newly funded in the last 2 years)**





# Focus on New Connections Across Disciplines

## AMO Coupled with PA and Gravitational Physics Very-Low Mass Dark Matter & Gravity at Small Distances

Blewitt and Derevianko: “Search of Topological Dark Matter with Atomic Clocks and GPS Constellation”: Analyze clock data from the 30 GPS satellites, which use atomic clocks for everyday navigation. Topological defect dark matter would cause initially synchronized clocks to become desynchronized, causing time discrepancies between spatially separated clocks to exhibit a distinct signature.



Geraci: “Measuring Gravity at the Micron Scale with Laser-Cooled Trapped Microspheres”: Uses levitated spheres of silica to search for hidden forces - sensitivities of a few zeptonewtons. Distinguishes between whether gravity will be much weaker than expected at short range, or stronger.





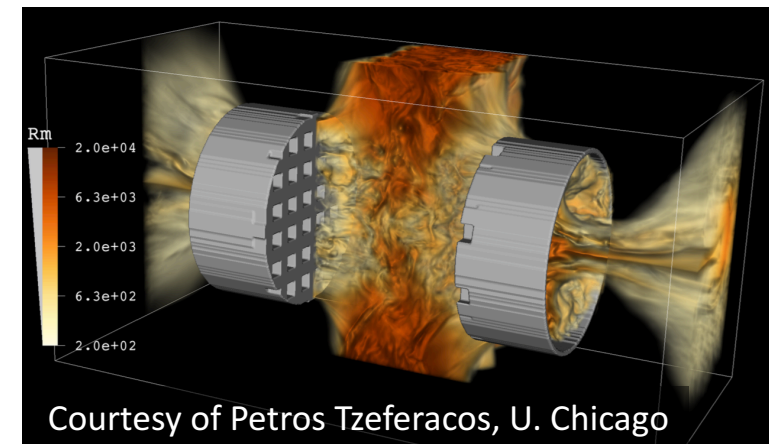
# Advances in Laboratory Plasma Astrophysics



Courtesy of Lawrence Livermore Nat'l Lab

## 3D MHD Simulations and High Power Laser Experiments Study the Turbulent Origin of Cosmic Magnetic Fields

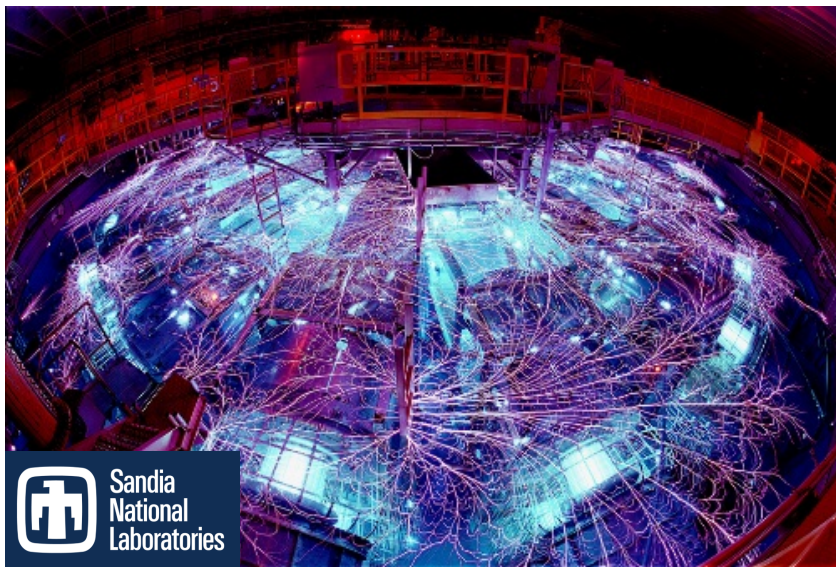
- A key astrophysical mechanism that is responsible for generation of the cosmic magnetic fields is being recreated in a terrestrial laboratory.
- Guided by three-dimensional numerical simulations with the FLASH code, experiments are being conducted in the largest laser facilities in the world – LLNL's National Ignition Facility and CEA's Megajoule Laser Facility – to demonstrate and characterize the turbulent dynamo mechanism.



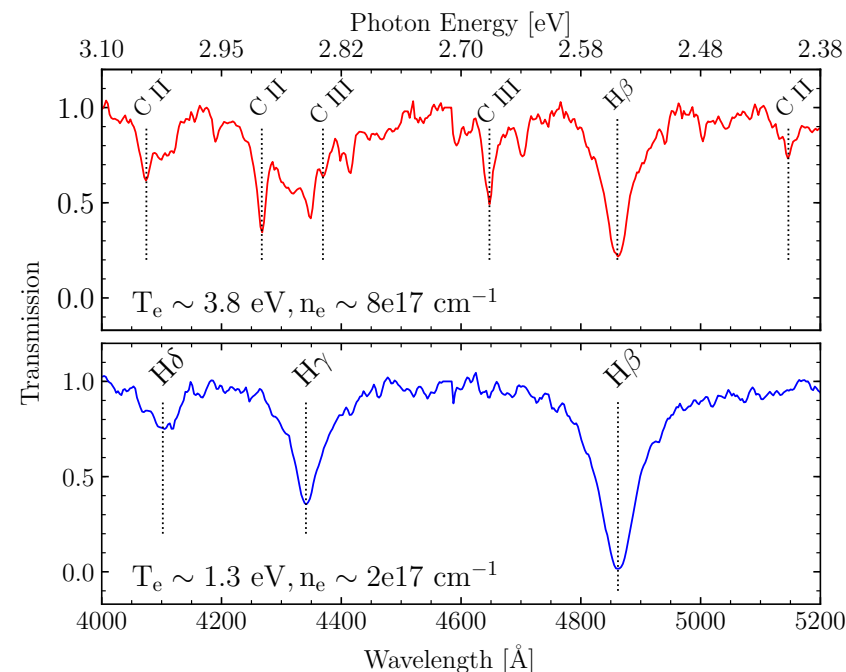
Mira BG/Q at Argonne Nat'l Lab



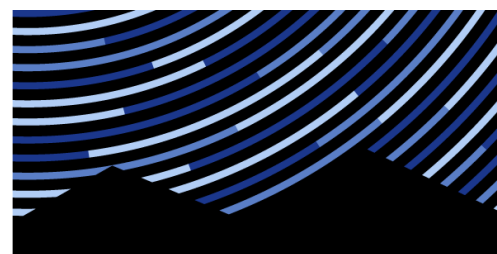
# Advances in Laboratory Plasma Astrophysics



## White Dwarf Photospheric Plasmas on the Z-machine at Sandia Nat'l Lab



- The plasmas have long path lengths, allowing measurements of simultaneous absorption and emission over a range of densities and temperatures characteristic of WD photospheres.
- Critical aspects of the atomic line-broadening theory of hydrogen, helium, and carbon are being examined. One goal is to improve the ability to diagnose the plasma conditions at the surface of WD stars to unlock their potential as cosmic chronometers.



**McDonald Observatory**  
THE UNIVERSITY OF TEXAS AT AUSTIN

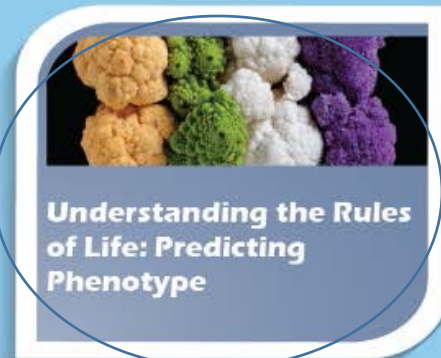
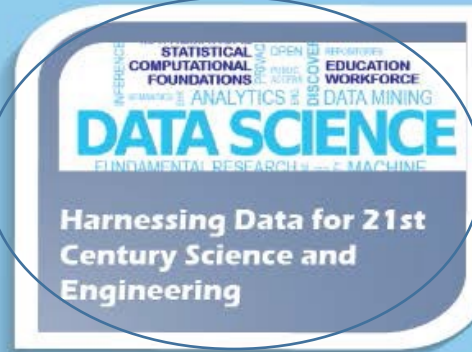


Courtesy of Don Winget, Center for Astrophysical Plasma Properties, University of Texas at Austin

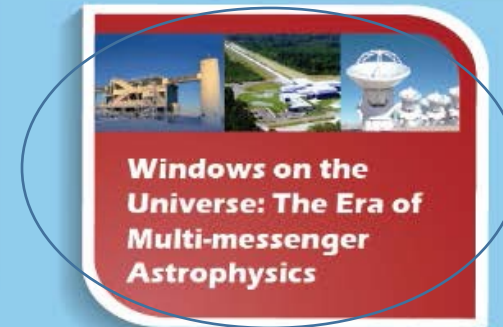




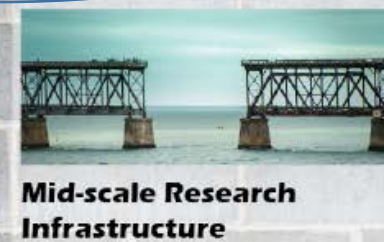
# Looking Ahead: Ten Big Ideas



## RESEARCH IDEAS



## PROCESS IDEAS





# Windows on the Universe

Multiple Forms of PHY Investment;  
Individual Investigators, Facilities, Centers;  
Partnerships guided by strong intellectual overlap

PHY Programs: Particle Astrophysics; Gravitational Physics;  
Nuclear Physics; Plasma Physics; Astrophysics &  
Cosmology Theory; Nuclear Theory

NSF Facilities: IceCube (GEO/PLR;MPS/PHY); NSCL (MPS/PHY)

Physics Frontiers Centers: NanoGrav (MPS/PHY/AST); JINA (MPS/PHY);  
KICP (MPS/PHY;GEO/PLR); KITP (MPS/PHY/AST/DMR;BIO/MCB)

Large Experiments: SPT, ACT, CMB (Especially appreciate CMB CDT), etc.

DOE Partnerships: LHC, SuperCDMS, HAWC, Plasma Partnership, etc.

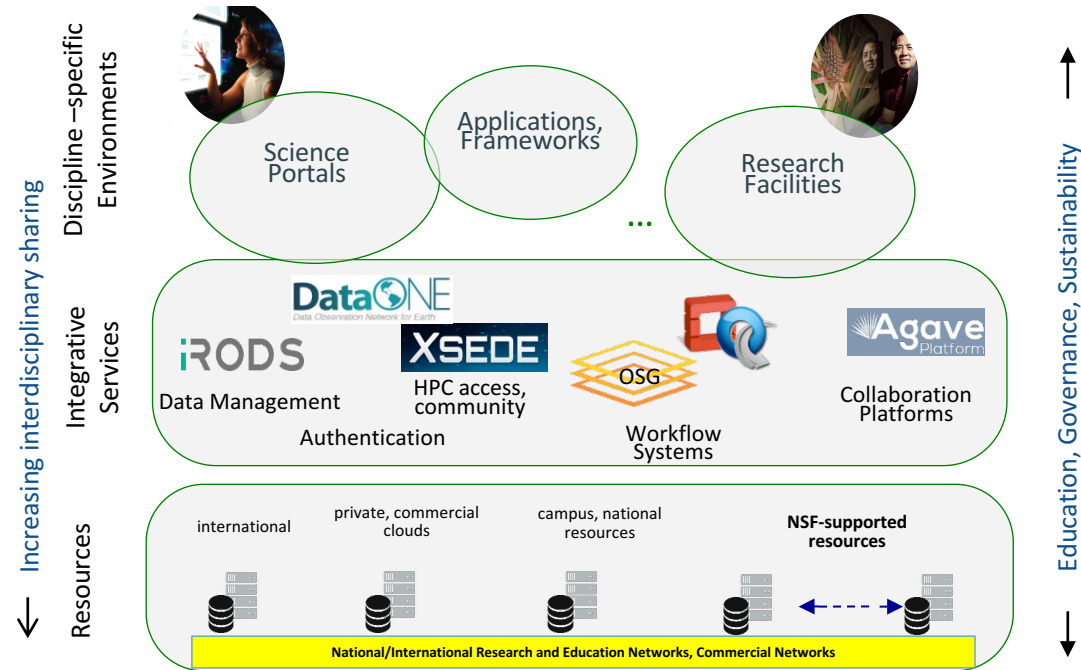
International: LHC, XENON1T, VIRGO, IceCube, etc.







# Harnessing Data



**Developing HDR cyberinfrastructure within a shared architectural vision**



# Harnessing Data

Domain Science  $\leftrightarrow$  Computation and Data Challenges

Connecting domain scientists with computer scientists

Essential for Handling Large Data Sets from LIGO, LHC, LSST

Efforts closely coordinated with OAC in CISE Directorate

Stay tuned for announcements of any special programs that offer possibilities for joint efforts related to Windows on the Universe