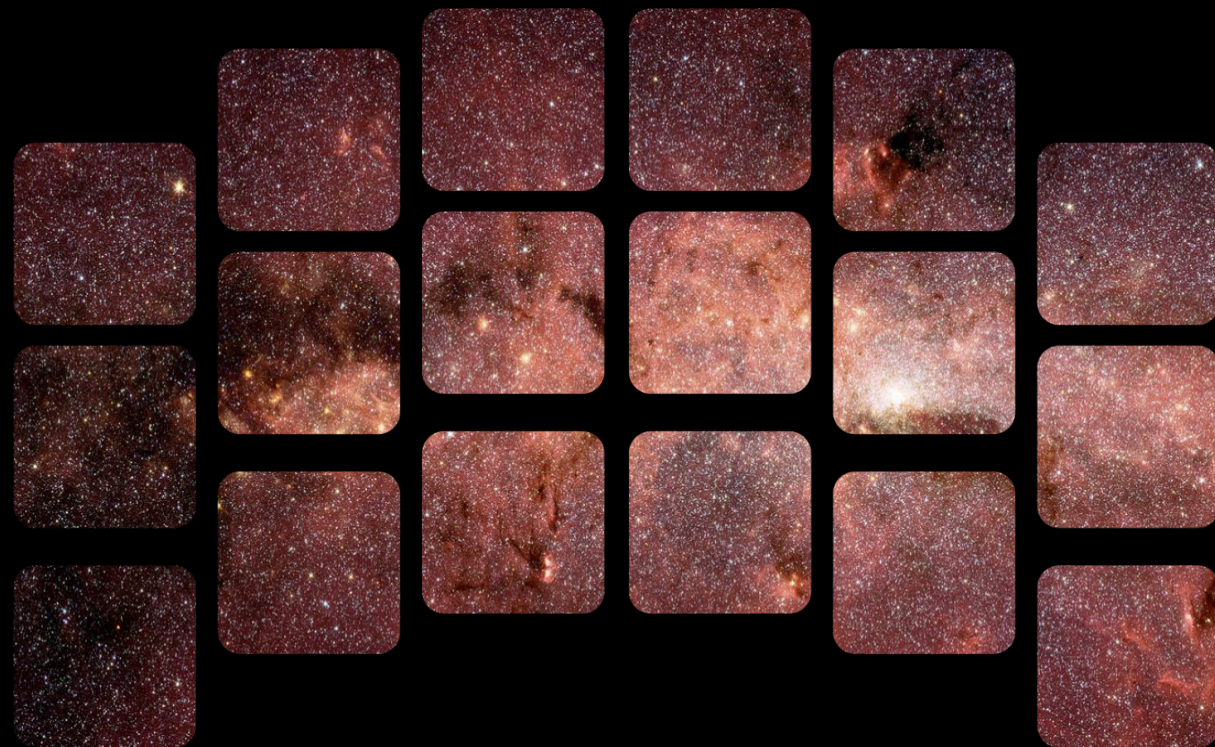


Project Status

R.ÖMAN



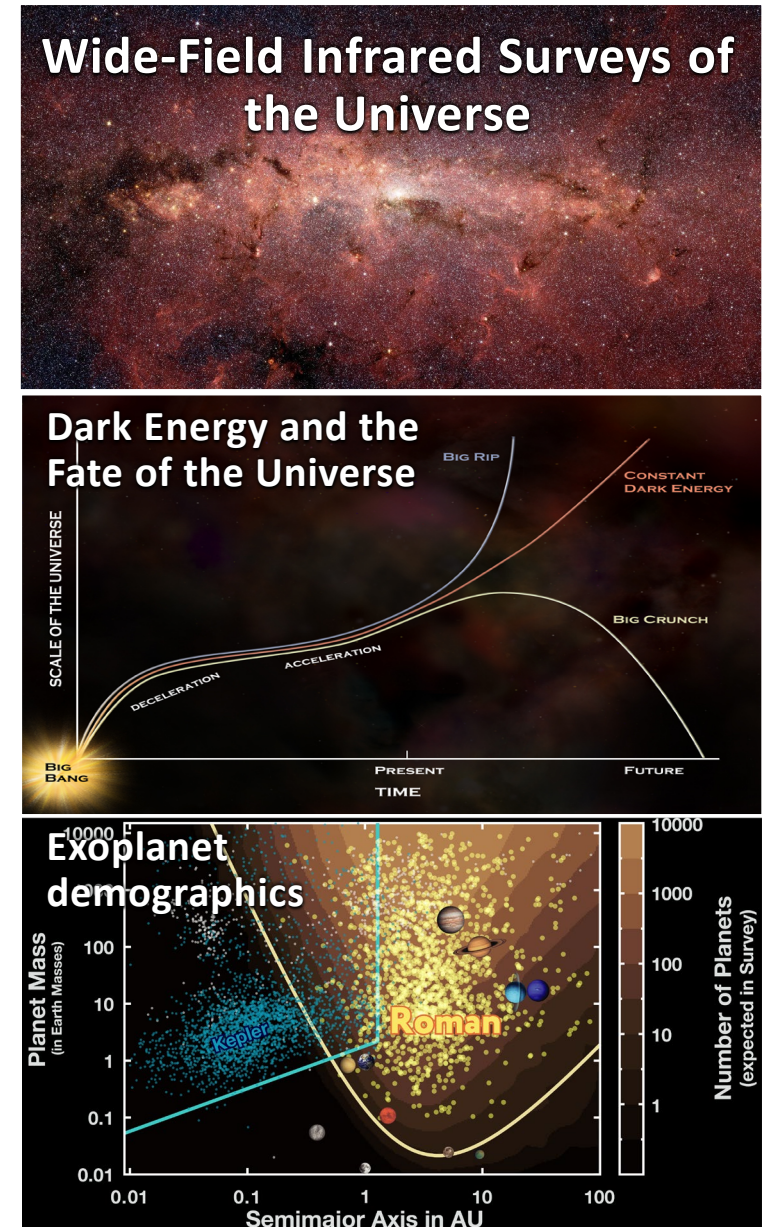
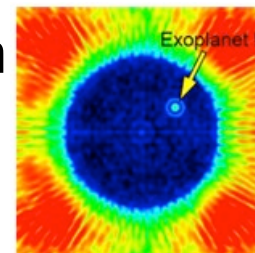
SPACE TELESCOPE

WFIRST is now the Nancy Grace Roman Space Telescope



Roman Mission Objectives

- **Wide Field Infrared survey**
 - Imaging and spectroscopy to >26.5 AB mag
- **Expansion history of the Universe**
 - Using supernova, weak lensing and galaxy redshift survey techniques
- **Growth of Structure in the Universe**
 - Weak lensing, redshift space distortions and galaxy cluster techniques
- **Exoplanet Census**
 - Statistical census of exoplanets from outer habitable zone to free floating planets
- **General Observer program**
 - Devote substantial fraction of mission lifetime to peer reviewed general observer program
- **Coronagraph technology demonstration**
 - Demonstrate exoplanet coronagraphy with active wavefront control

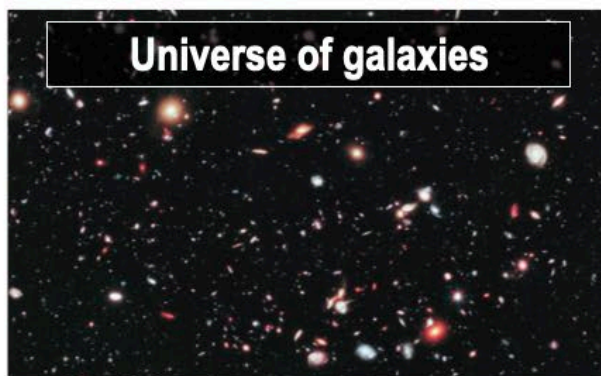


Roman's very Broad Science Menu includes:

Evolution of the Universe



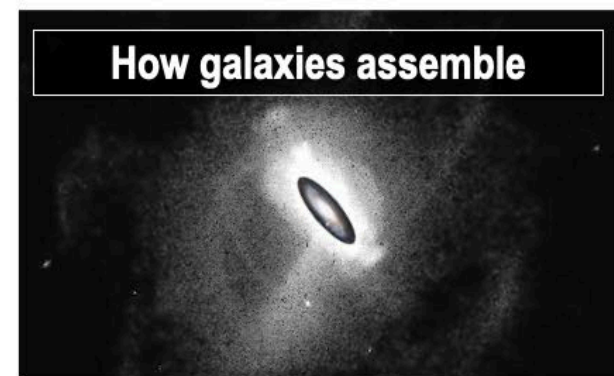
Universe of galaxies



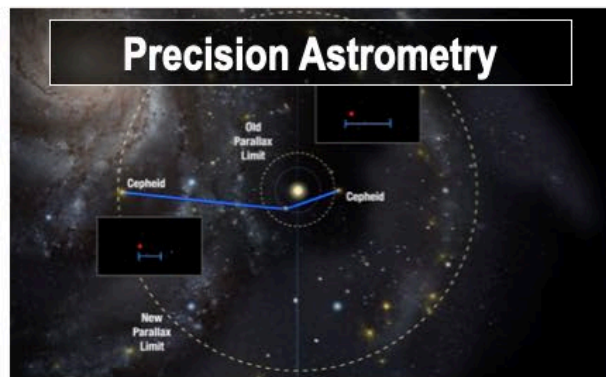
Mapping dark matter



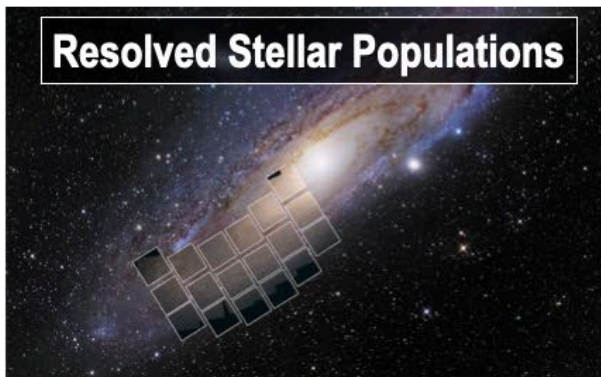
How galaxies assemble



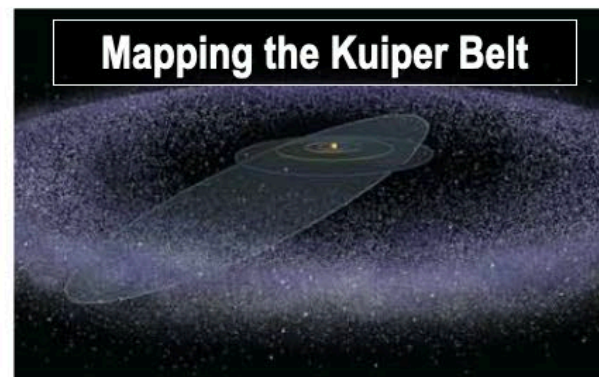
Precision Astrometry



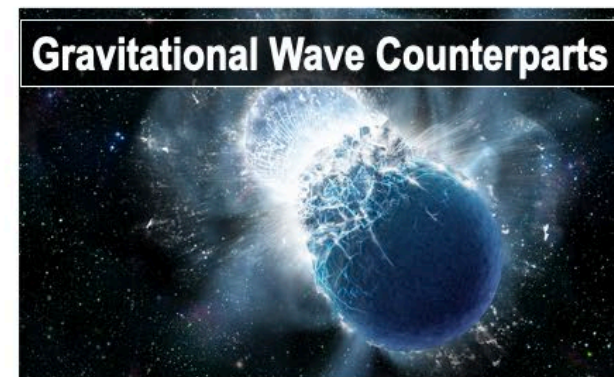
Resolved Stellar Populations



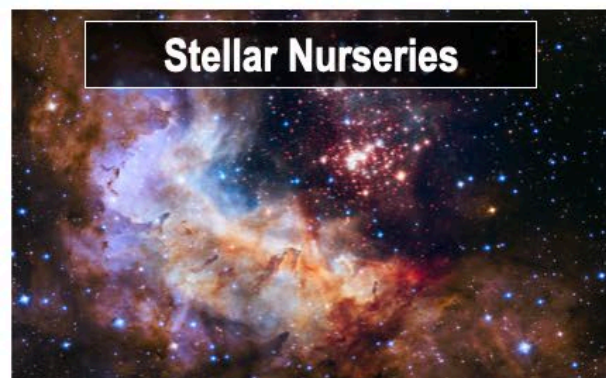
Mapping the Kuiper Belt



Gravitational Wave Counterparts



Stellar Nurseries



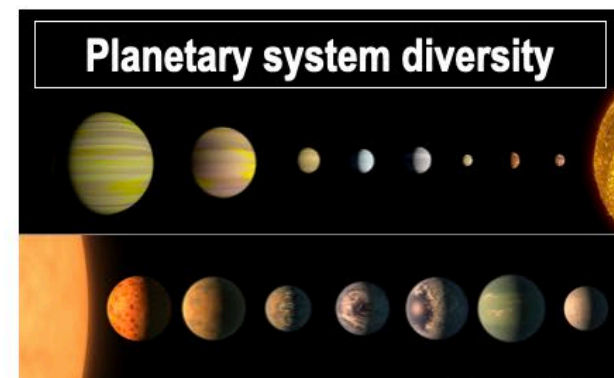
Asteroseismology



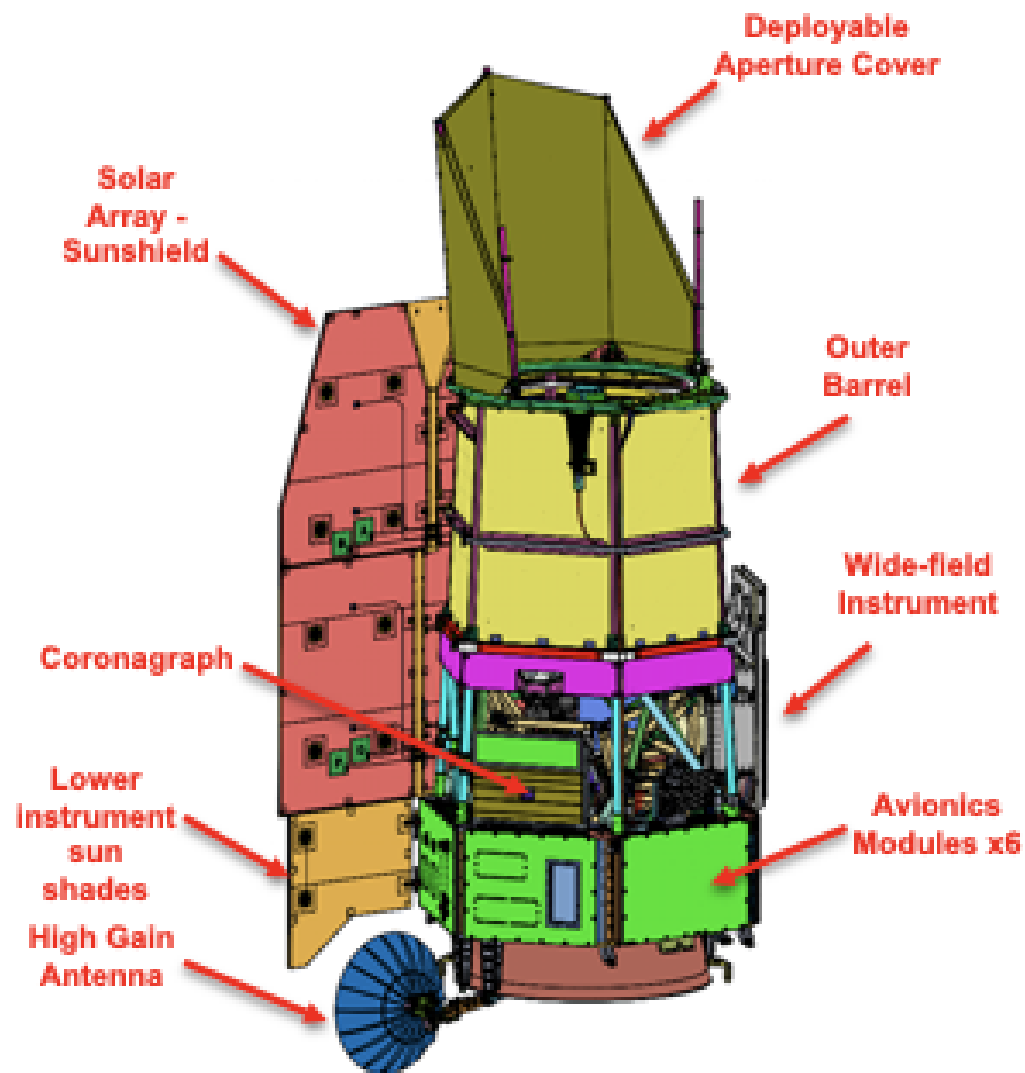
Exoplanet Direct Imaging



Planetary system diversity



Roman Observatory and Instruments



Telescope: 2.4m aperture

Two Instruments:

Wide Field Imager / Slitless Spectrometer

- Vis/Near IR bandpasses (0.48 – 2.3 micron)
- Field of view 0.281 deg² (~200× HST WFC3-IR)
- 18 4k × 4k detectors (288 Mpixels)

Coronagraph

- Visible bandpass
- Contrast 10⁻⁸-10⁻⁹

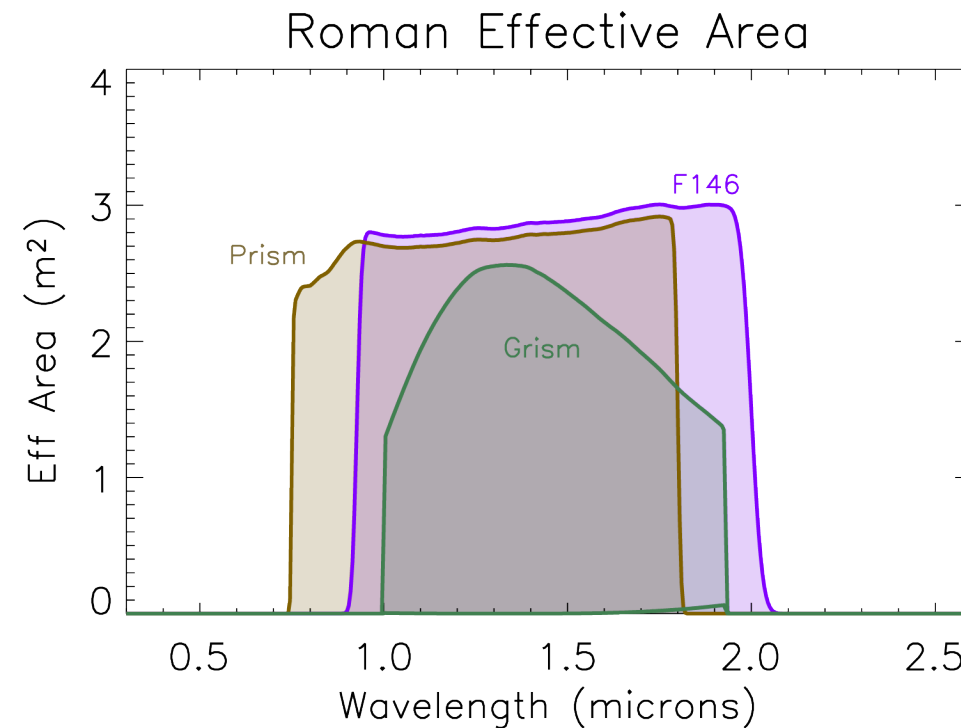
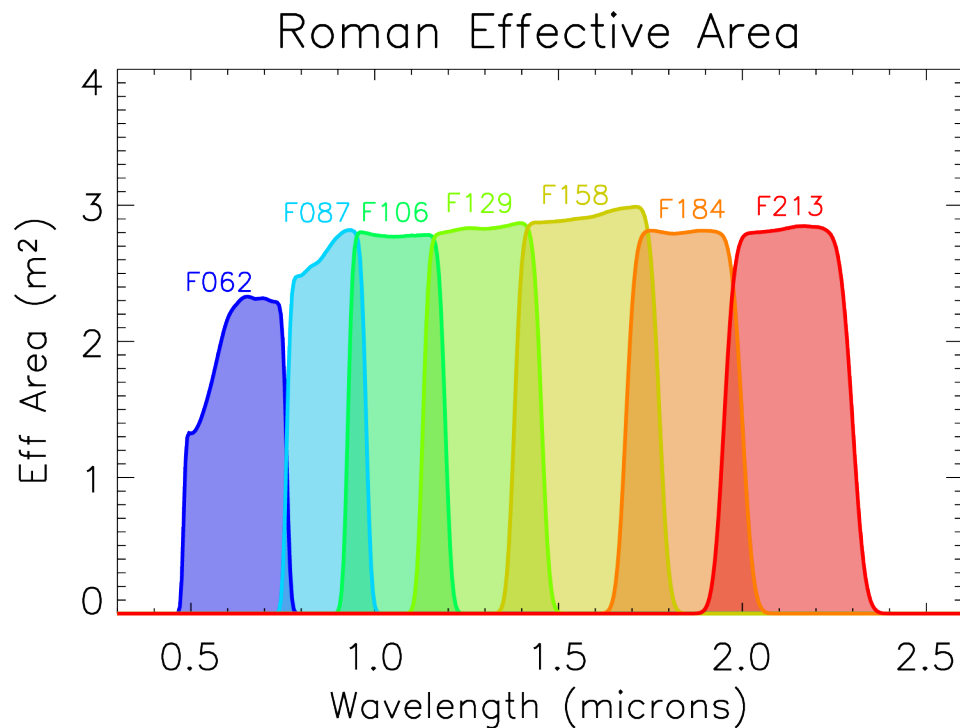
Data Volume: 11 Tb/day

Orbit: Sun-Earth L2

Mission Duration: 5 yr, 10yr goal

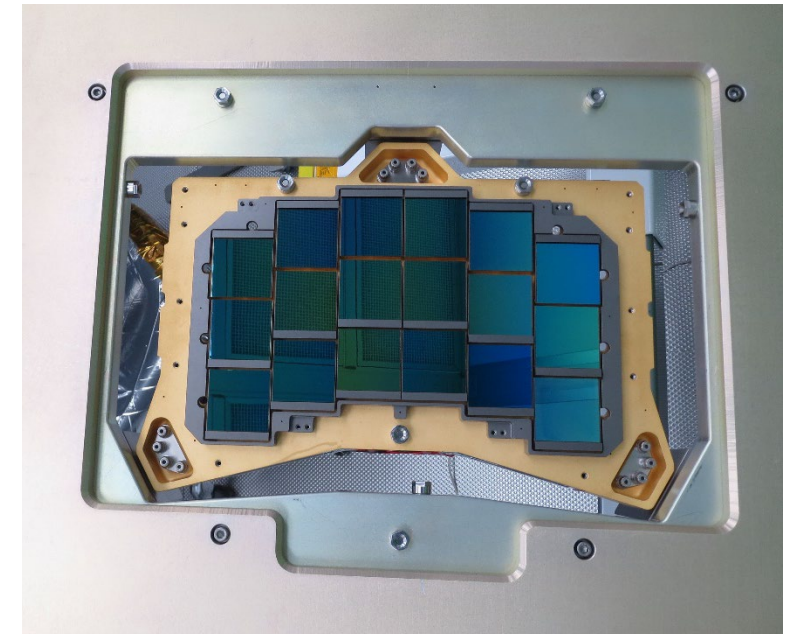
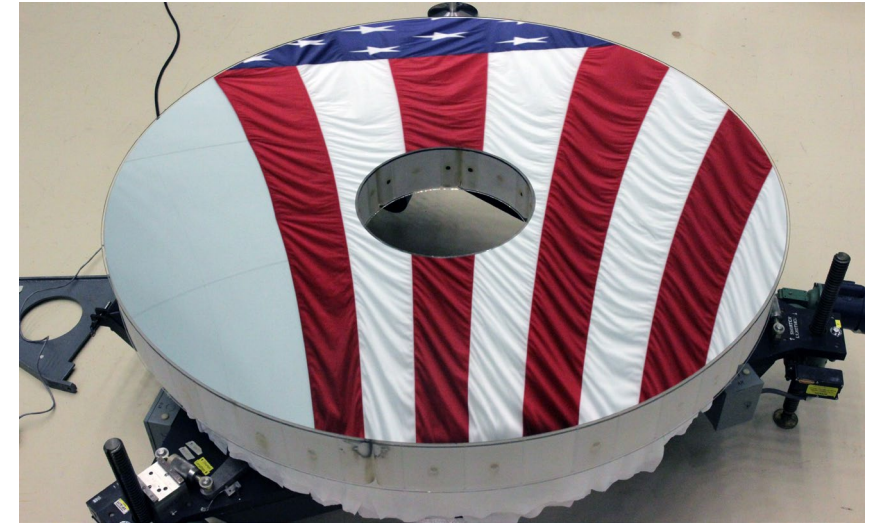
Wide Field Instrument: Filters/Dispersers

- Imaging capability is ~ 1 sq. deg. per day at any band to 5σ AB depth of F062=29.6, F087=29.2, F106=29.1, F129=29.1, F158=29.1, F184=28.5, F213=27.2, F146=29.4.
- Spectroscopy via prism (0.6-1.8 μ m, $R\sim 100$, ~ 24 AB) and grism (1.0-1.9 μ m, $R\sim 600$, ~ 22 AB)



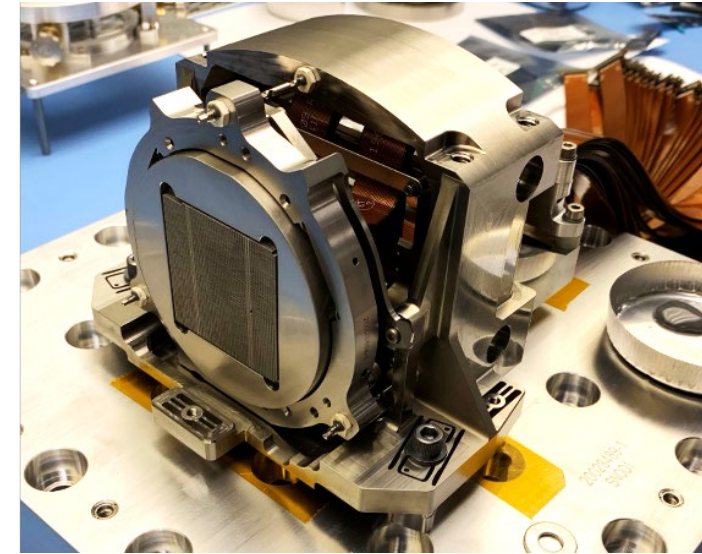
Roman Space Hardware Telescope Status

- **Advanced to phase C (i.e. started implementation phase) in February 2020**
- **Mission Critical Design Review in Sept 2021**
- **Flight hardware being built**
 - **Telescope (L3Harris):** Primary and secondary mirrors have been refigured, polished and coated; coronagraph relay optics polished and coated;
 - **Wide Field Instrument (GSFC/Ball):** Completed installation and alignment of all 18 engineering test unit (ETU) sensor chip assemblies (SCA) on the ETU mosaic plate
 - 15 out of 18 flight candidate SCAs in hand
 - added new F213 filter (1.95-2.3 micron) – now have imaging filters covering entire spectral range supported by mirrors/detectors!



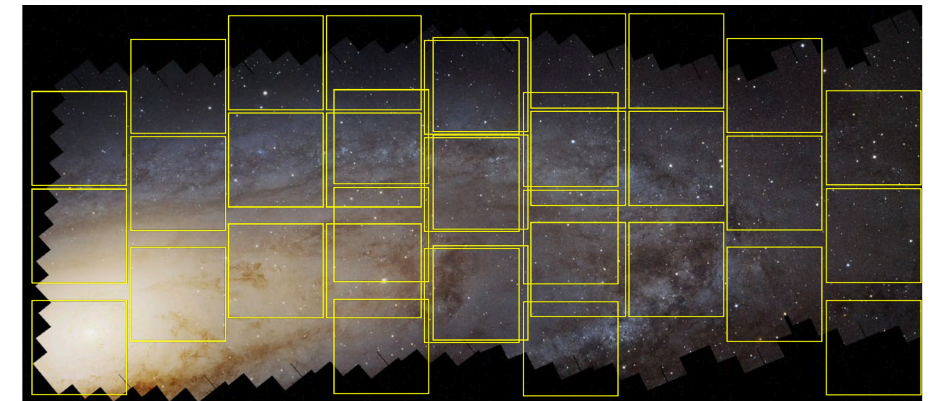
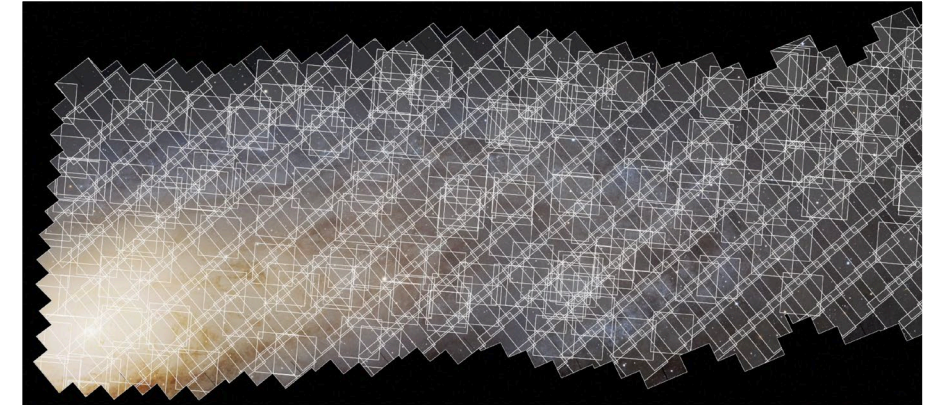
Roman Space Telescope Hardware Status cont.

- **Coronagraph (JPL)**
 - Deformable mirror technology has been demonstrated
 - Two flight actuators in hand
- **Spacecraft (GSFC)**
 - Procurement of flight subsystems well underway
 - Mechanical Hardware Engineering Development Units (EDUs) nearly complete
- **On track for launch in mid-2020's**



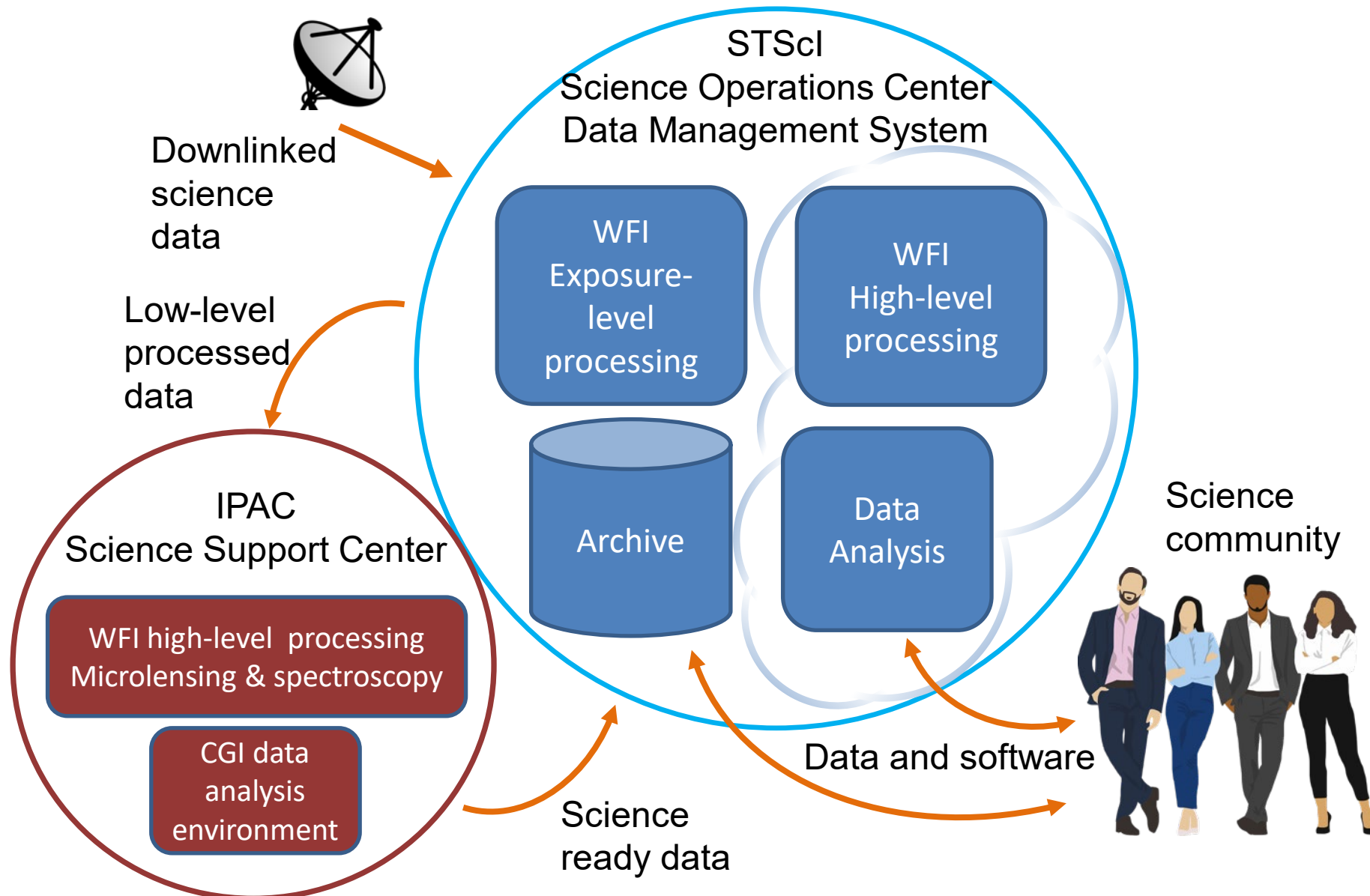
Roman as a Precise Survey Facility

- The power of Roman is not *just* that it has a large field of view:
 - Very efficient observations
 - Rapid slew & settle
 - no Earth occultations
 - no South Atlantic Anomaly
 - Well understood and stable PSF
 - Stable thermal environment (L2 orbit, thermal control of all parts of the optical system)
 - Rigid optical structure with vibration isolation from the spacecraft
 - Stable attitude control
 - Excellent flux calibration
 - Relative calibration system



For details, see Akeson et al. 2019 <https://arxiv.org/abs/1902.05569>

Roman Science Data Management



Roman Science Operations

- **Data system consists of:**
 - Pipeline for low level data processing
 - Pipeline for high level processing
 - Science platform (HLPP) allowing users to interact with data and high-level processing software in the cloud
 - Archive with HST/JWST/MAST like functionality
- **Continuing to refine data processing plans at STScI (SOC) and IPAC (SSC) on relative roles between science centers and science teams**
 - Expanded catalog functions to be implemented by science centers: deblending, photo-z, some time-domain functions etc
 - Astrometry functions – e.g. tying to *Gaia*
 - Some elements of PSF characterization
 - Instrument simulations
- **Evaluating science platform and cloud computing options and solutions**
 - Including considerations of interoperability with other facilities (e.g. *Rubin*)

Science Investigation Teams

- **Supernova Cosmology: Ryan Foley, Saul Perlmutter**
- **Nearby Galaxies: Ben Williams**
- **Extragalactic: Brant Robertson**
- **Weak Lensing and Galaxy Redshift Survey: Olivier Dore**
- **Exoplanet Coronagraphy: Bruce Macintosh, Margaret Turnbull**
- **Archival Research: Alexander Szalay**
- **Cosmic Dawn: James Rhoads**
- **Exoplanet Microlensing: Scott Gaudi**
- **Milkyway: Jason Tumlinson**

- **~300 scientists in total**
 - scientific performance requirements related to the specific science area,
 - design of overall observational strategy concept,
 - science data analysis techniques,
 - ground and space calibration requirements,
 - science simulations, precursor observations,
 - ground calibration, observational needs, data processing, ancillary data collection/incorporation, analysis, dissemination and documentation of the proposed science investigation.

- **Current science team contracts expire later this year**

Adjutant Scientists
David Spergel - WFI
Jeremy Kasdin - CGI

Roman Science Interest Group

- <https://roman.gsfc.nasa.gov/science/rsig.html>
- **Meeting presentations and notes available on the meetings tab**
 - Recent discussions have been on the observing program
- **Annual opportunities to join this group**
- (for SOC role at STScI) see also: <https://www.stsci.edu/roman/about/roman-advisory-committee-rstac>

Megan Donohue (Chair)	Michigan State U.
Zeljko Ivesic	U. Washington
Jessica Lu	UC Berkeley
John MacKenty	STScI
Ashley Villar	Columbia U / Flatiron Institute
Alice Shapley	UCLA
Keith Bechtol	UW, Madison
Saurabh Jha	Rutgers U
Peter Melchior	Princeton U
Dara Norman	NOIRlab
Jessie Christiansen	NEXSci/ CalTech
Rachel Bean	Cornell U
Ryan Hickox	Dartmouth
Lisa Storrie-Lombardi	Las Cumbres Observatory
Dimitri Mawet	CalTech

Roman Opportunities

- **Opportunities for participation in *Roman* offered in ROSES-2021**
 - **Key Project Teams:** Science teams to conduct scientific investigations using the data from the major surveys identified by the Astro2010 Decadal Survey
 - **Coronagraph Community Participation Program:** Investigators to work with the coronagraph instrument team to plan and execute tech demo observations
 - **Wide Field Instrument Preparatory Science:** Investigators to work on science preparation activities related to mission performance verification and science operations preparation

- **All *Roman* observing time is available through open processes**
 - Major surveys will be defined using a community-driven open process
 - Key Projects – funded science investigations using these surveys – openly competed
 - Roman observing time will be available for General Observer (GO) projects
 - All data will be available to the community with no period of limited access

Roman Observations

- **Core Community Surveys:** a significant fraction of the prime mission used for revolutionary surveys of unprecedented scale
- **Three Core Community Surveys that provide broad scientific power and address 2010 Decadal Survey science goals**
 - Extragalactic Wide Area Survey
 - Extragalactic Time Domain Survey
 - Galactic Time Domain Survey
- **The definition of Core Community Surveys will be established via an open process, with a goal of maximizing the overall science return while simultaneously meeting the cosmology and exoplanet science requirements**
- **Several calls for General Observer surveys closer to, and after, launch**
- **No proprietary period for any *Roman* data**

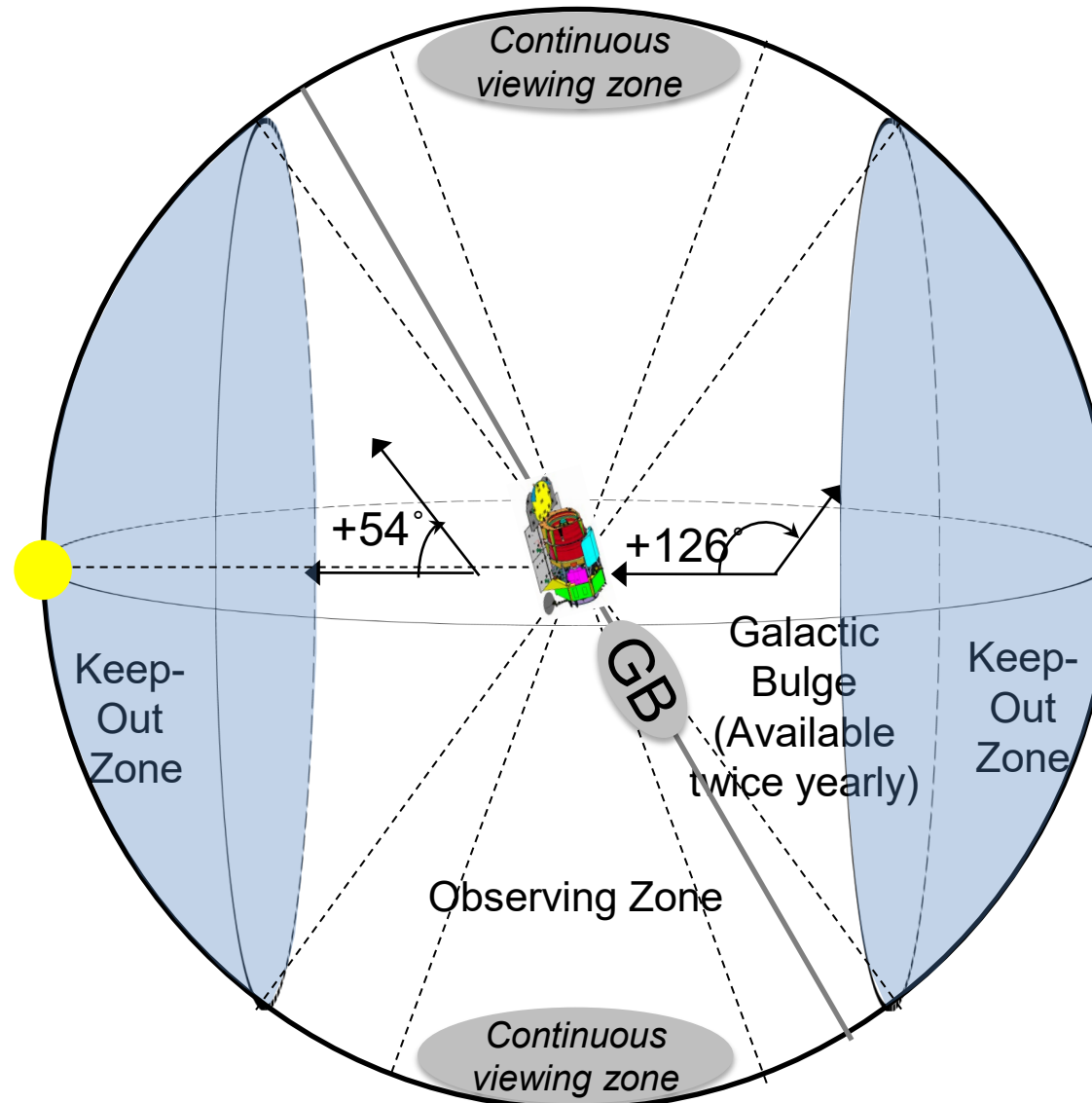
Roman Field of Regard

Observing Zone:

- 54° - 126° off Sun Line
- 360° about Sun Line
- $\pm 15^\circ$ about line of sight

Slew/settle times are rapid – typically ~1 min for adjacent field of view

Earth/Moon avoidance angles are a minor sporadic constraint



Extragalactic Time Domain Survey in fields within 20° of the ecliptic poles, located in continuous viewing zone(s)

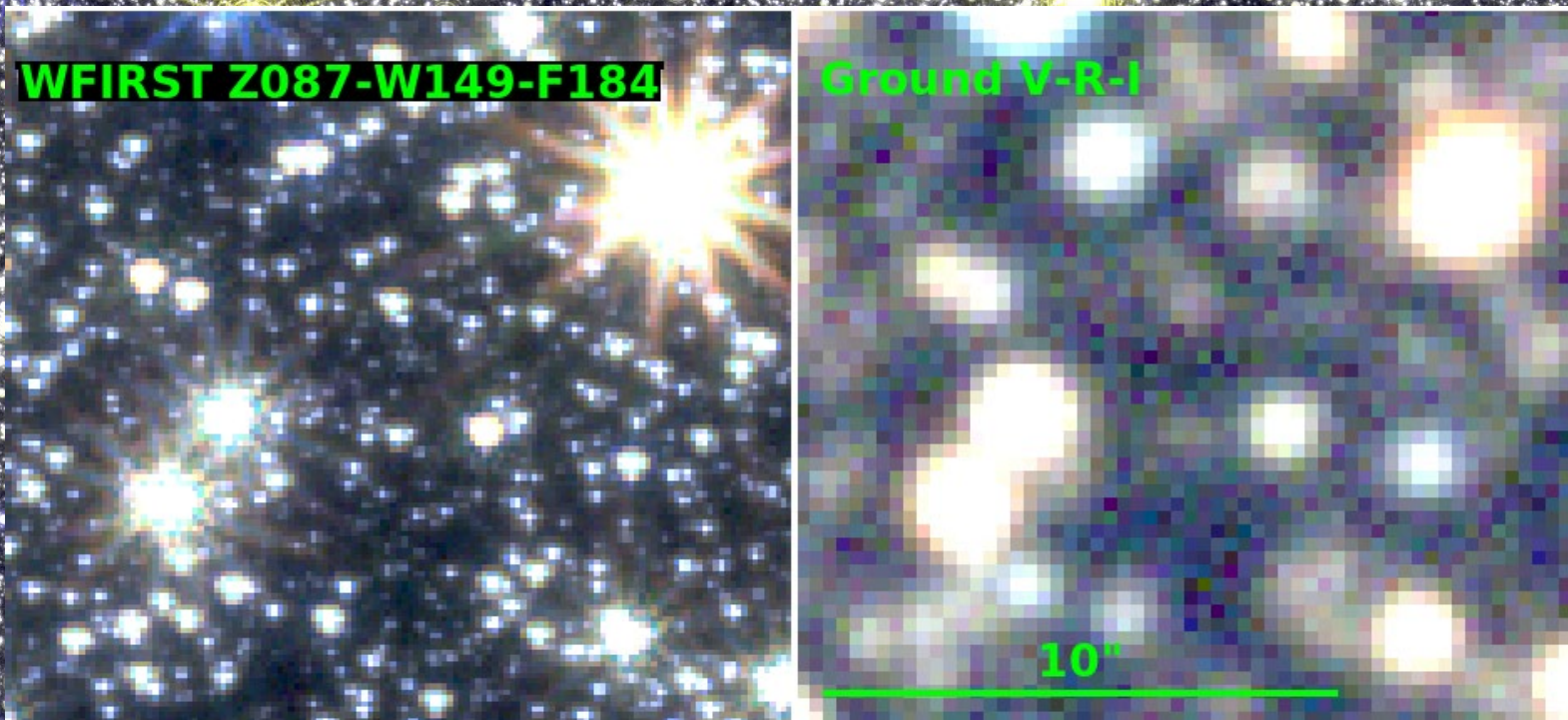
Extragalactic Wide Area Survey (also GO & Coronagraph) observations can be located within the full Observing Zone

Galactic Time Domain Survey can observe inertially fixed fields in the Galactic Bulge (GB) for 72 days twice a year

Roman Galactic Time Domain Survey

- *Roman's* time domain imaging survey of the Galactic Bulge
- Envisioned as ~2 square degrees in wideband infrared imaging every 15 minutes, plus other band(s) every ~day
- Over 40,000 images across 5 years = ~2Ms
- Depth ~25AB (epoch), ~31AB (coadded)
- Detects hundreds of millions of stars in each epoch
- Enables exoplanet demographics via microlensing
- But this is just the beginning...
 - precise photometry & stellar parameters for ~100M stars
 - asteroseismology of millions of red giants
 - millions of distances from parallax
 - a hundred thousand transiting planets!
 - thousands of trans-Neptunian objects

~1% of *Roman*'s field



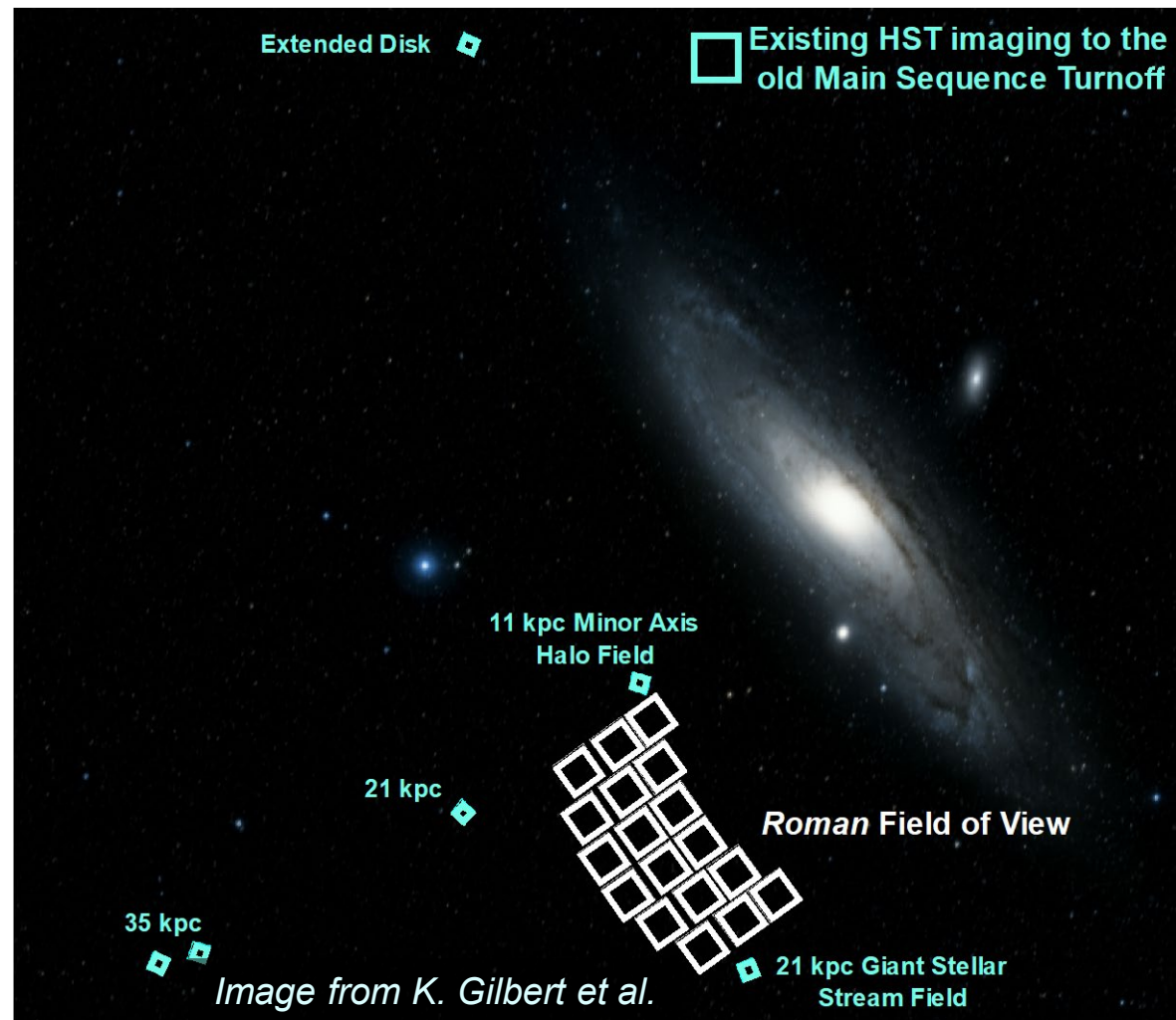
Roman General Observer Surveys

- *Roman* will devote ≥ 400 days of telescope time in the first five years to surveys defined by GO opportunities
- Envisioned as a few tens of projects of significant scale, substantial science outcomes, and lasting value to the community

Studies of absolutely anything are permitted for submission; taking galaxy assembly as an example...

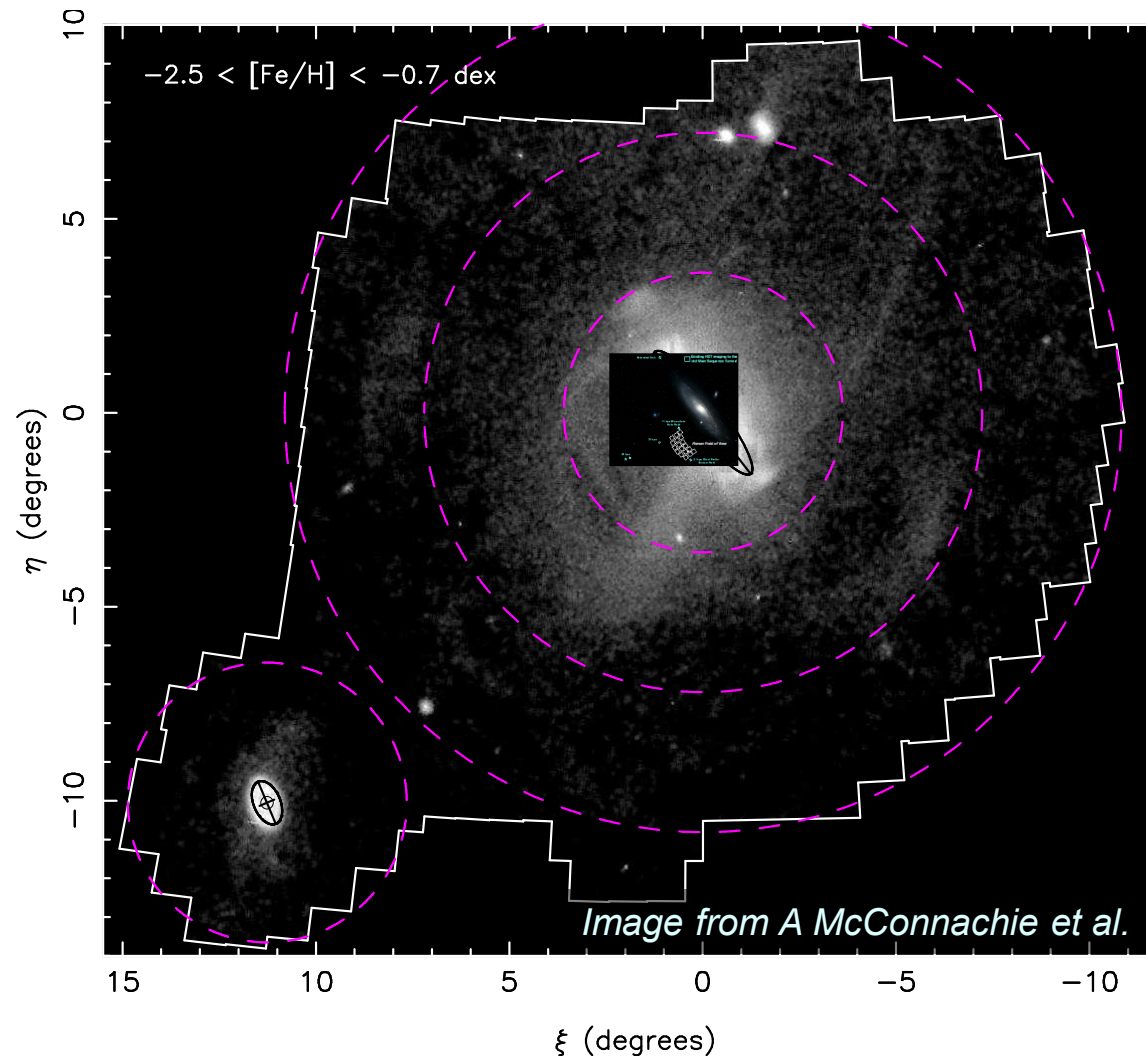
Roman General Observer Surveys

Studies of absolutely anything are permitted for submission; taking galaxy assembly as an example...



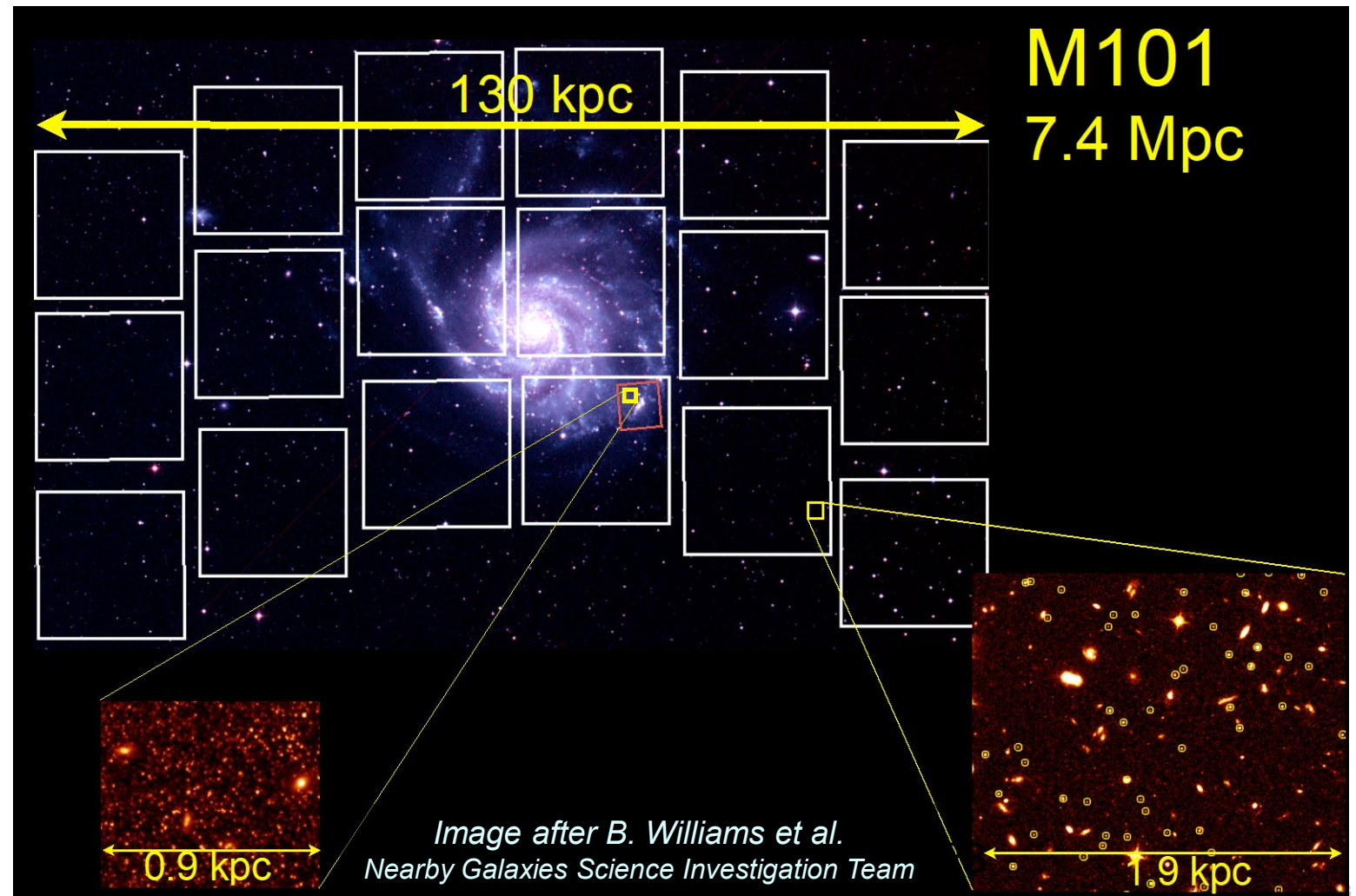
Roman General Observer Surveys

Studies of absolutely anything are permitted for submission; taking galaxy assembly as an example...



Roman General Observer Surveys

Studies of absolutely anything are permitted for submission; taking galaxy assembly as an example...

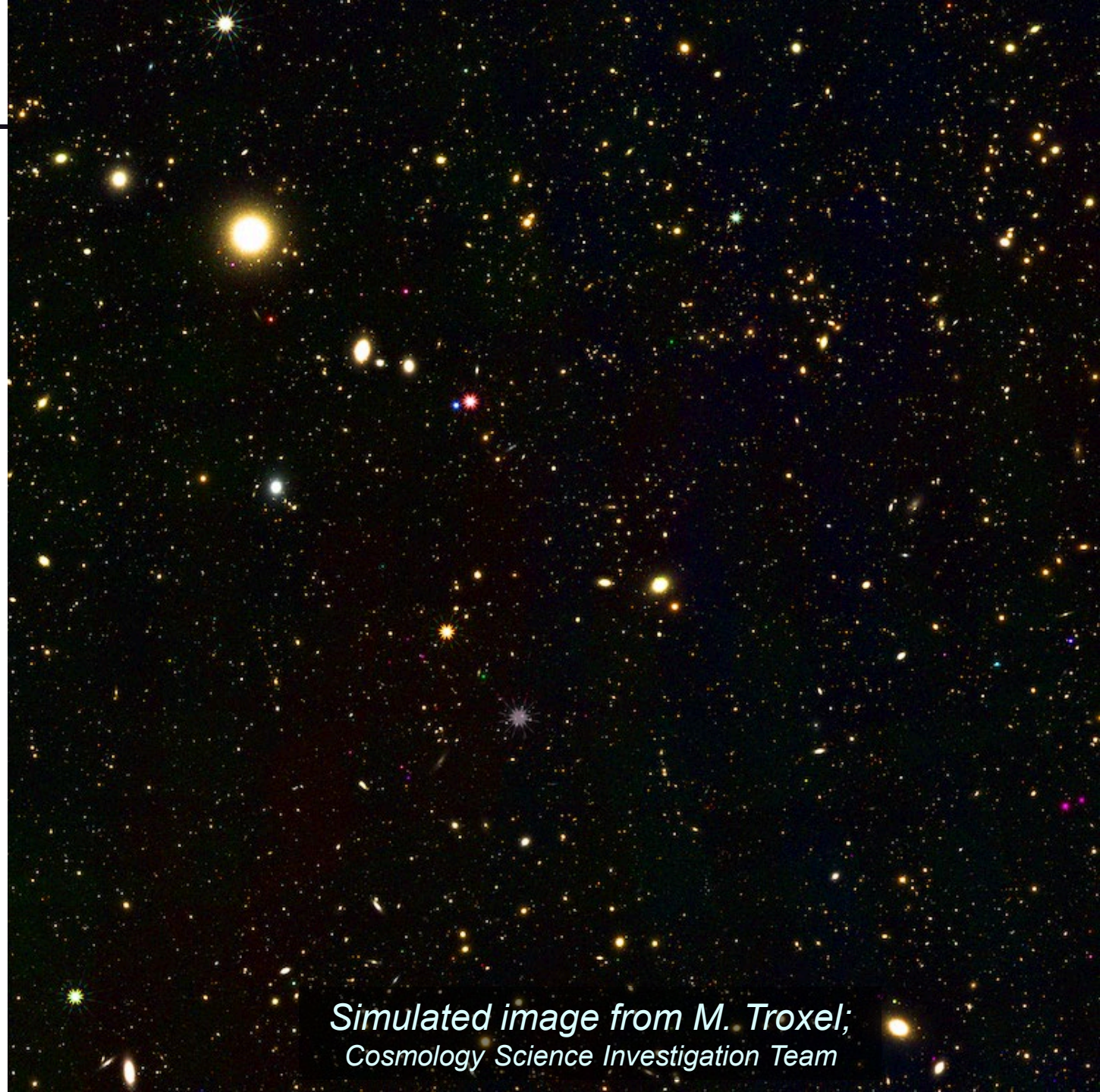


Roman Extragalactic Wide Area Survey

- *Roman's* wide+deep imaging and spectroscopic survey
- Envisioned as $\sim 2,000$ square degrees in multiband infrared imaging & slitless spectroscopy at $\sim 1\text{-}2\mu\text{m}$
- Depth $\sim 27\text{AB} / 10^{-16} \text{ erg/s/cm}^2$ (5σ)
- Detects hundreds of millions of galaxies (imaging) and tens of millions (spectroscopically), mostly at redshifts 1-3 but with significant numbers beyond $z=10$!
- Enables cosmology techniques such as weak lensing & baryon acoustic oscillations
- But this is just the beginning...
 - tens of thousands of galaxy clusters
 - thousands of galaxies at redshift 10-15
 - stellar streams and dwarfs around nearby galaxies

Roman Imaging

- YJH
- ~27AB
- 1/18th FOV
- ~30min



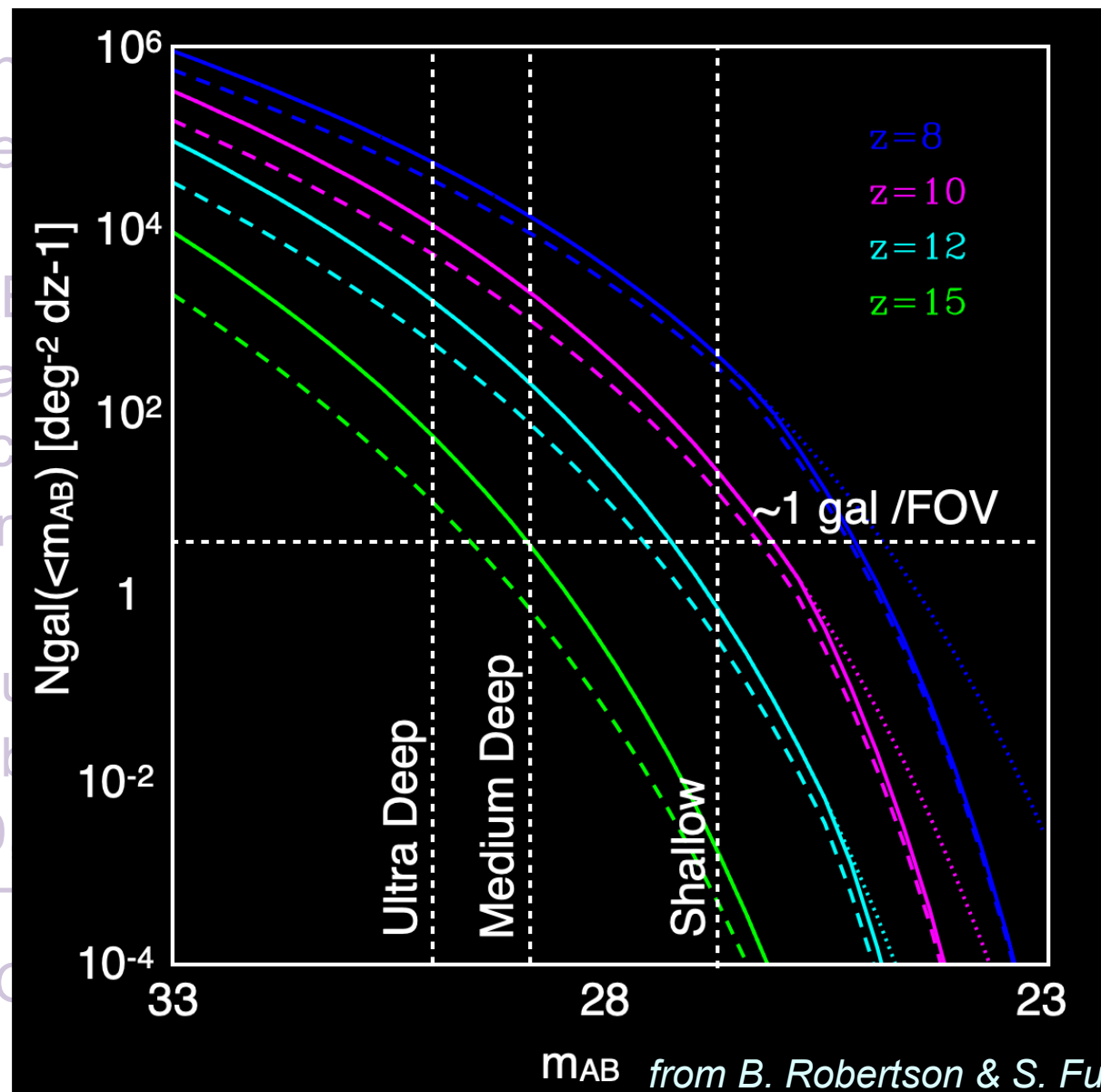
*Simulated image from M. Troxel;
Cosmology Science Investigation Team*

Roman Extragalactic Time Domain Survey

- *Roman's* tiered, temporal imaging and spectroscopic survey
- Envisioned as ~ 20 square degrees in five-band infrared imaging & slitless spectroscopy at $\sim 0.6\text{-}2\mu\text{m}$
- Depth $\sim 25\text{-}26\text{AB}$ (epoch), $\sim 30\text{AB}$ (coadded)
- Track the brightness of more than a million galaxies every ~ 5 days for ~ 2 years with unprecedented photometric accuracy
- Enables supernova cosmology measurements
- But this is just the beginning...
 - variability of active galactic nuclei
 - proper motion $\sim 10\mu\text{as}$; possible extragalactic parallax
 - ~ 3000 Ia supernova, but $\sim 10,000$ supernovae total
 - transients (e.g. TDEs; SPRITEs?)
 - distances from RR Lyrae; tRGB

Roman Extragalactic Time Domain Survey

- *Roman's* tiered, temporal imaging
- Envisioned as ~20 square degree spectroscopy at ~0.6-2 μ m
- Depth ~25-26AB (epoch), ~30AB
- Track the brightness of more than 10⁶ galaxies with unprecedented photometric precision
- Enables supernova cosmology
- But this is just the beginning...
 - variability of active galactic nuclei
 - proper motion ~10 μ as; possible astrometry
 - ~3000 Ia supernova, but ~10⁶ transients (e.g. TDEs; SPRIT)
 - distances from RR Lyrae; tRC



m_{AB} from B. Robertson & S. Furlanetto

Roman Extragalactic Science Investigation Team

Engaging the Community

- *All Roman* observing time is available through open processes
- Major **Core Community Surveys** will be defined using a community-driven *open process*
- Observing time available via **General Observer** *proposal opportunities*
- **Key Projects** – funded science investigations using these surveys – will be *openly competed*
- **Preparatory Science** – funded work preparing for any of the above, also via *proposal opportunities*
- All data will be available to the community with *no period of limited access*